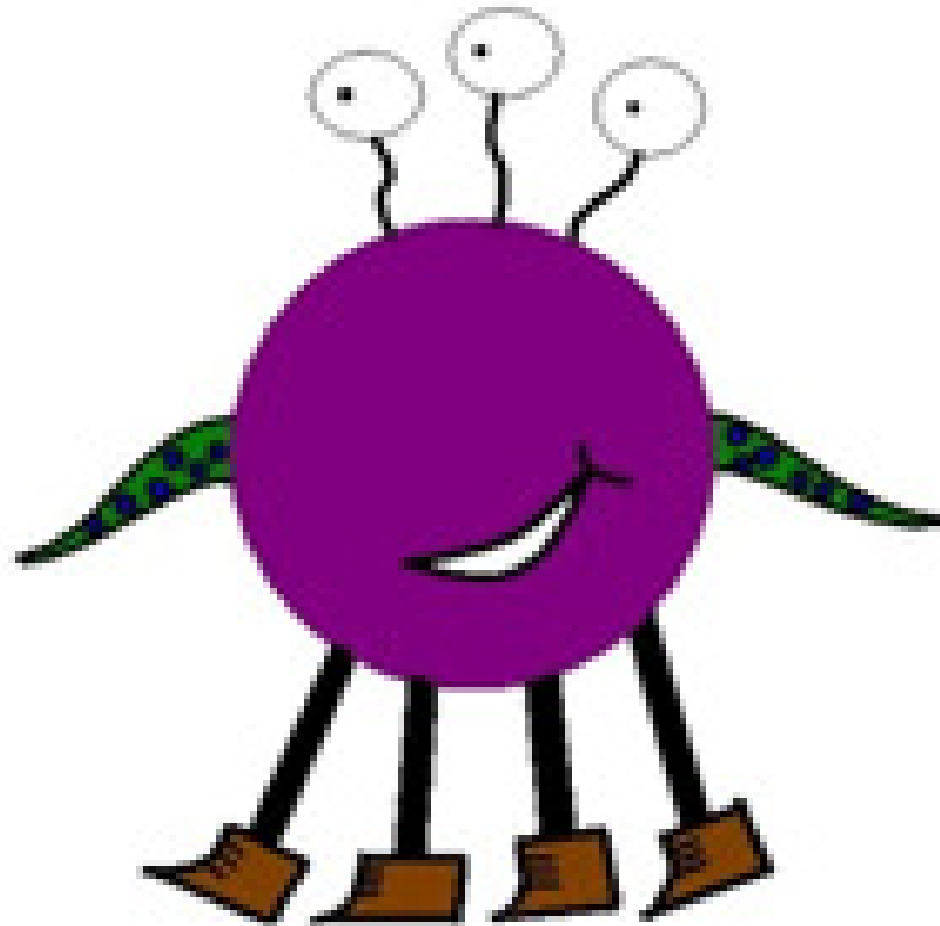


Exotic Physics in High Energy Cosmic Rays

Bernardo Tomé
7th NWAP09, S. Tomé



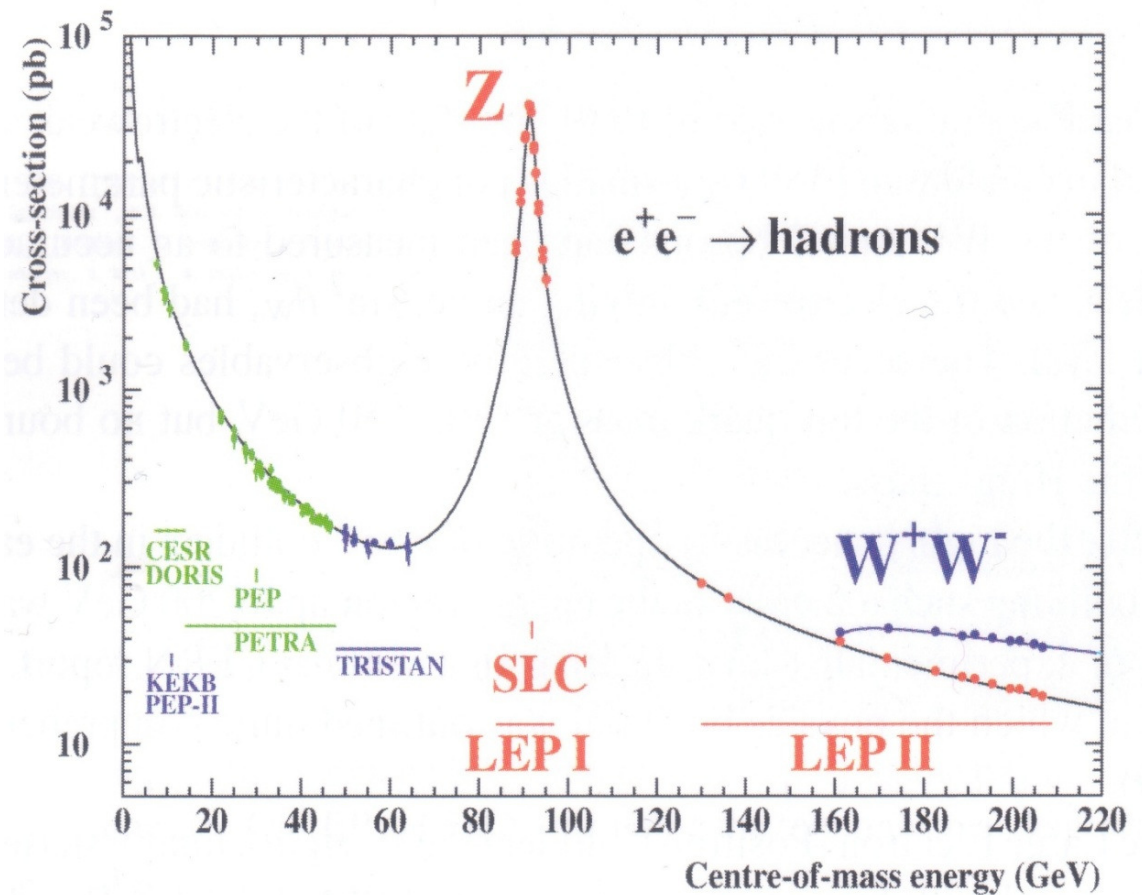
This talk is largely inspired in several talks presented at the [1st Pierre Auger Workshop on Exotic Physics in Cosmic Rays](#) held at Lisboa last July.

Acknowledgements to M.C.Espírito Santo, I.Albuquerque, Washington Jr., D. Schuster

The success of the SM...

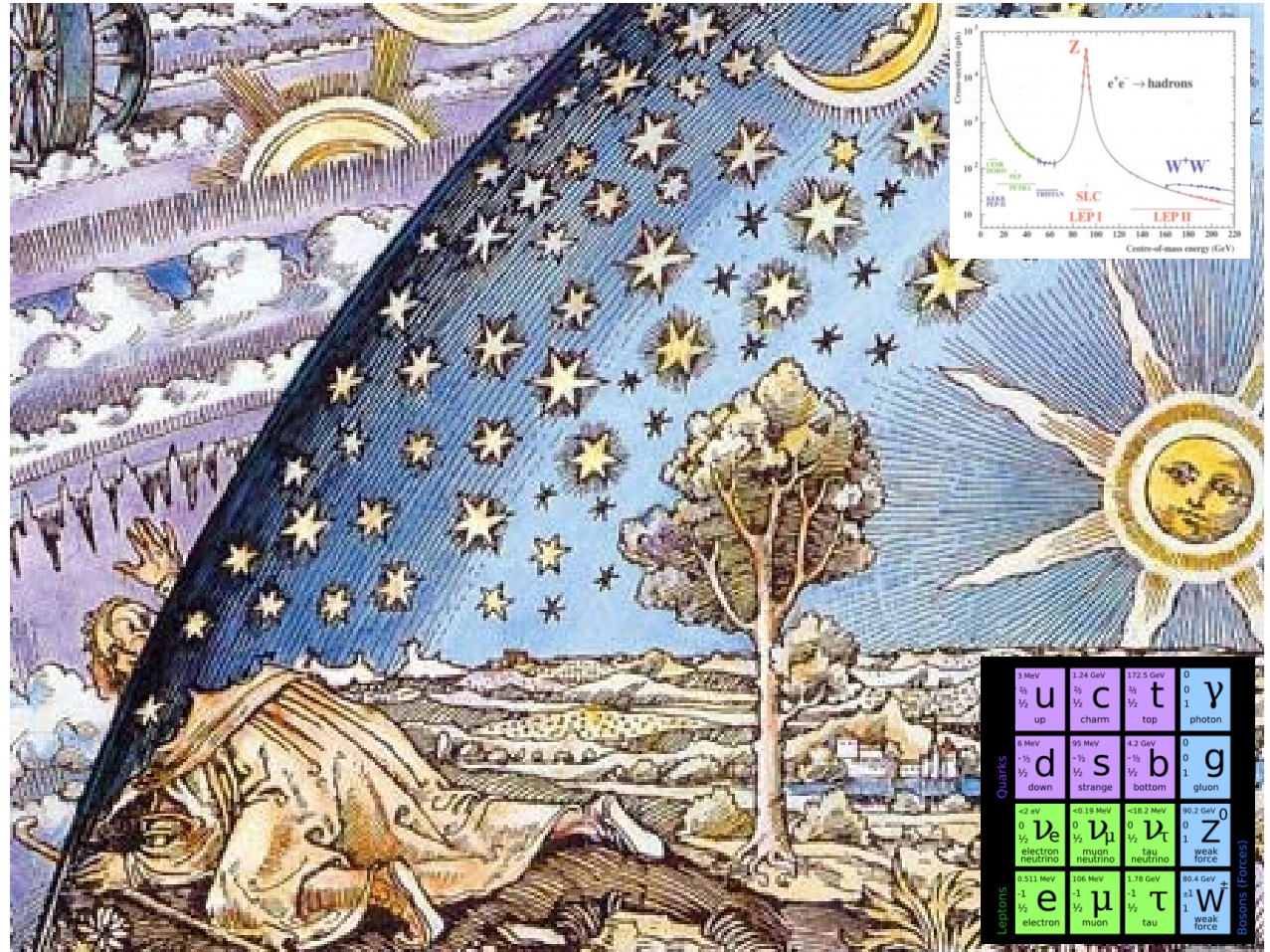
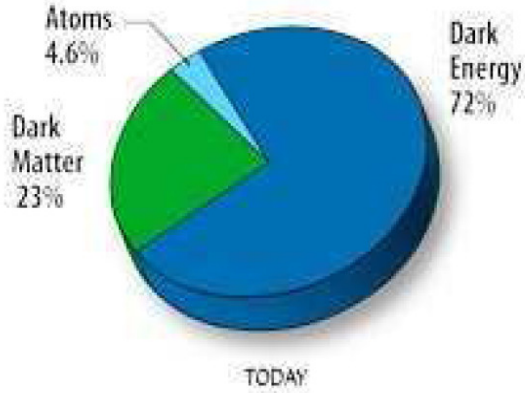
| | | | | |
|---------|---|---|---|--|
| Quarks | 3 MeV $\frac{2}{3}$ $\frac{1}{2}$ u up | 1.24 GeV $\frac{2}{3}$ $\frac{1}{2}$ c charm | 172.5 GeV $\frac{2}{3}$ $\frac{1}{2}$ t top | 0 0 1 γ photon |
| | 6 MeV $-\frac{1}{3}$ $\frac{1}{2}$ d down | 95 MeV $-\frac{1}{3}$ $\frac{1}{2}$ s strange | 4.2 GeV $-\frac{1}{3}$ $\frac{1}{2}$ b bottom | 0 0 1 g gluon |
| Leptons | <2 eV 0 $\frac{1}{2}$ ν_e electron neutrino | <0.19 MeV 0 $\frac{1}{2}$ ν_μ muon neutrino | <18.2 MeV 0 $\frac{1}{2}$ ν_τ tau neutrino | 90.2 GeV 0 1 Z weak force |
| | 0.511 MeV -1 $\frac{1}{2}$ e electron | 106 MeV -1 $\frac{1}{2}$ μ muon | 1.78 GeV -1 $\frac{1}{2}$ τ tau | 80.4 GeV ± 1 1 W [±] weak force |

Bosons (Forces)

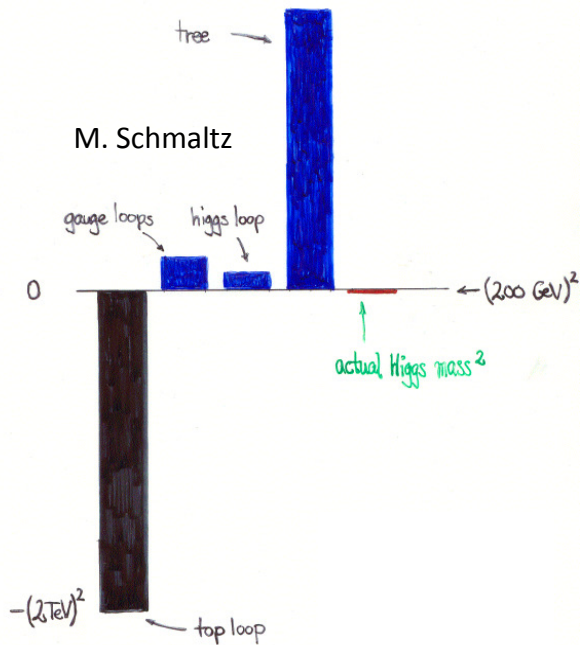


The need for New Physics !

95% of the Universe not explained by SM of particle physics

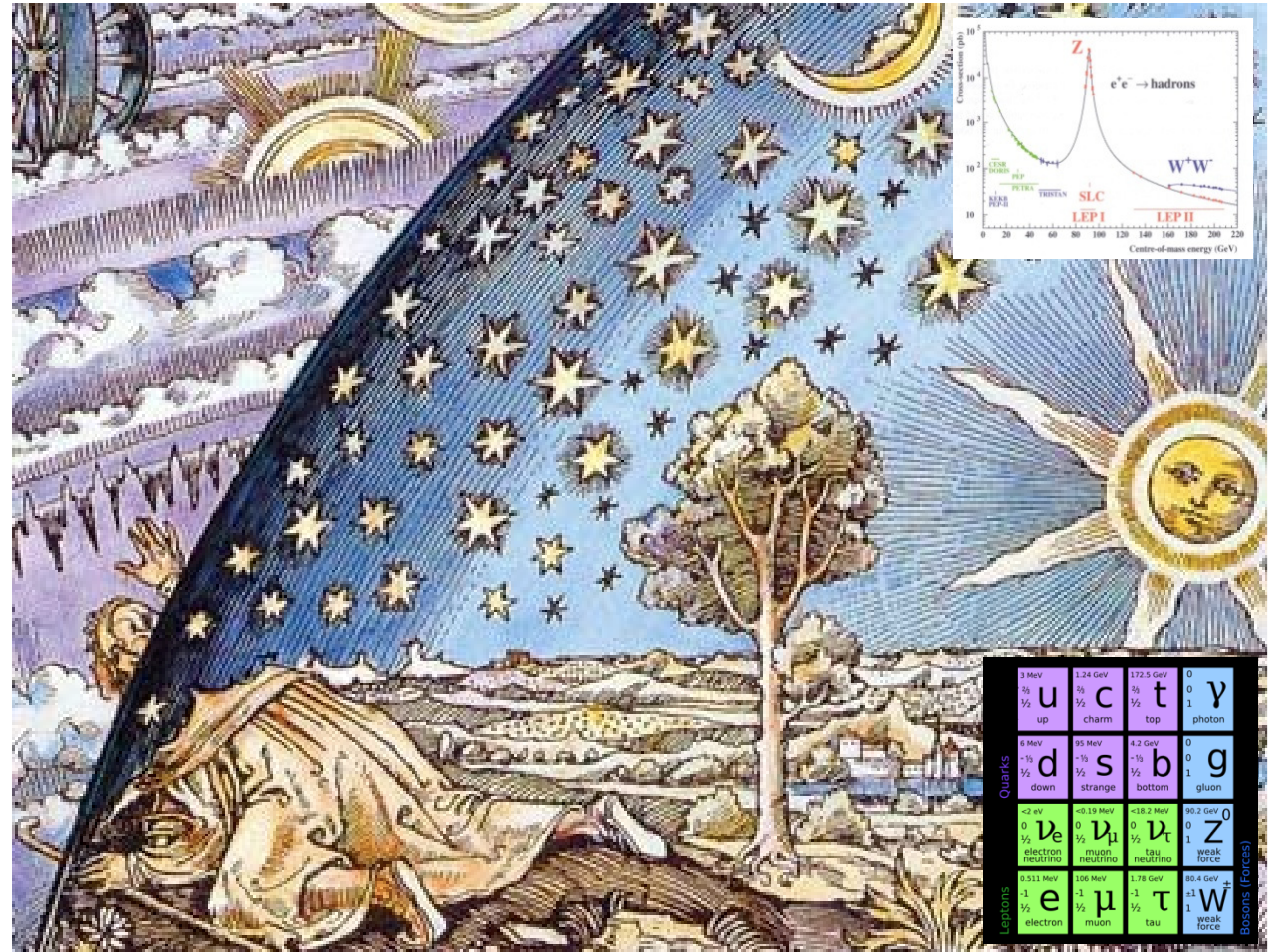
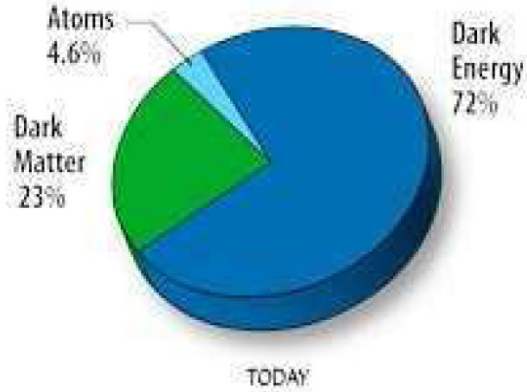


Fine tuning the Higgs mass : YUCK !!!

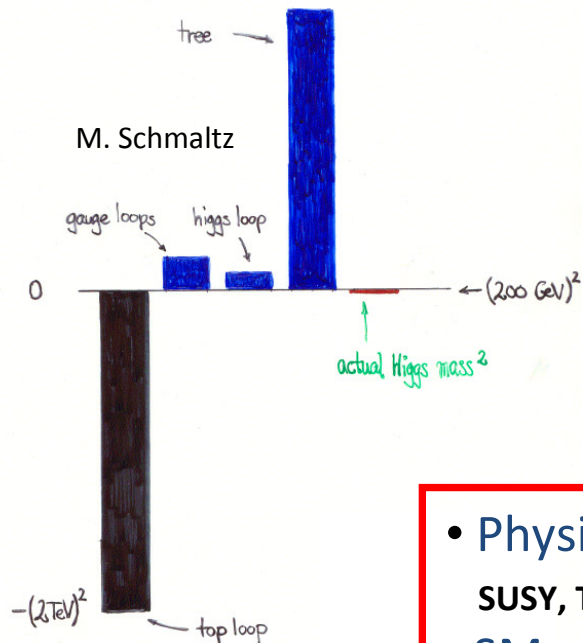


The need for New Physics !

95% of the Universe not explained by SM of particle physics

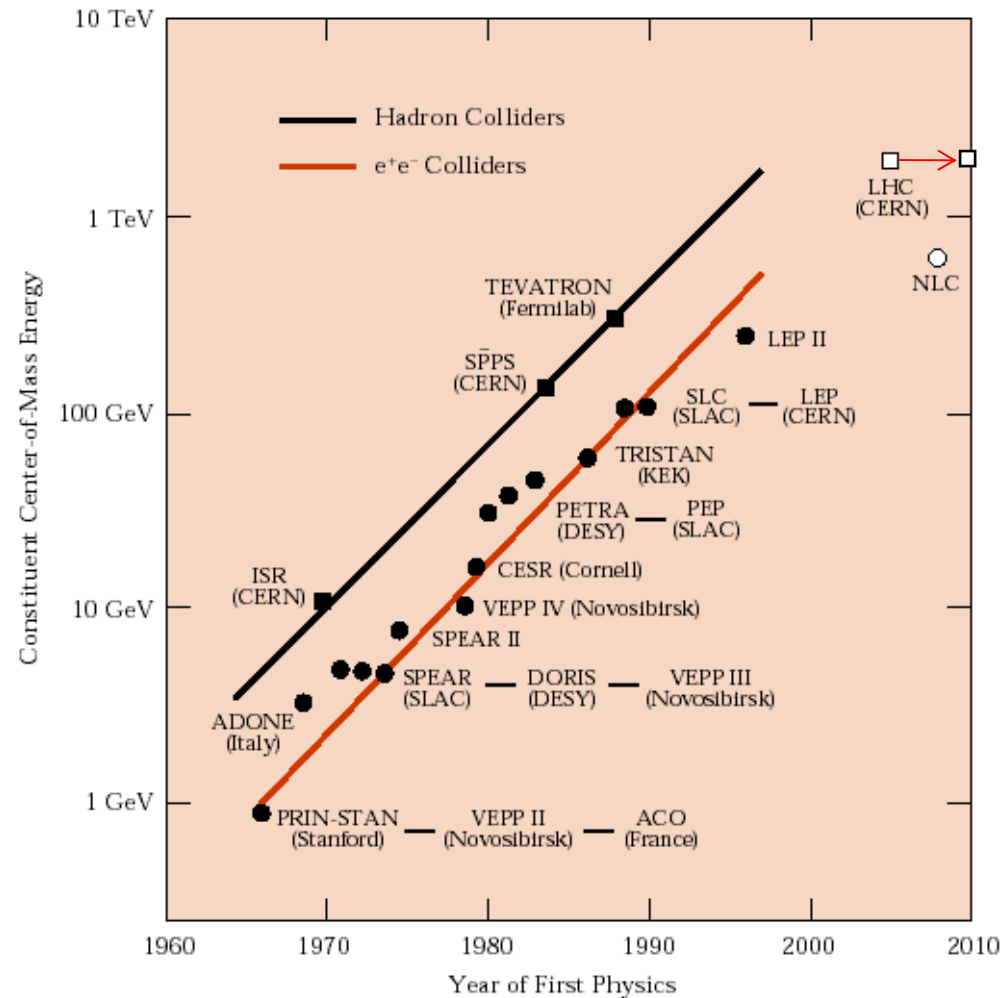


Fine tuning the Higgs mass : YUCK !!!

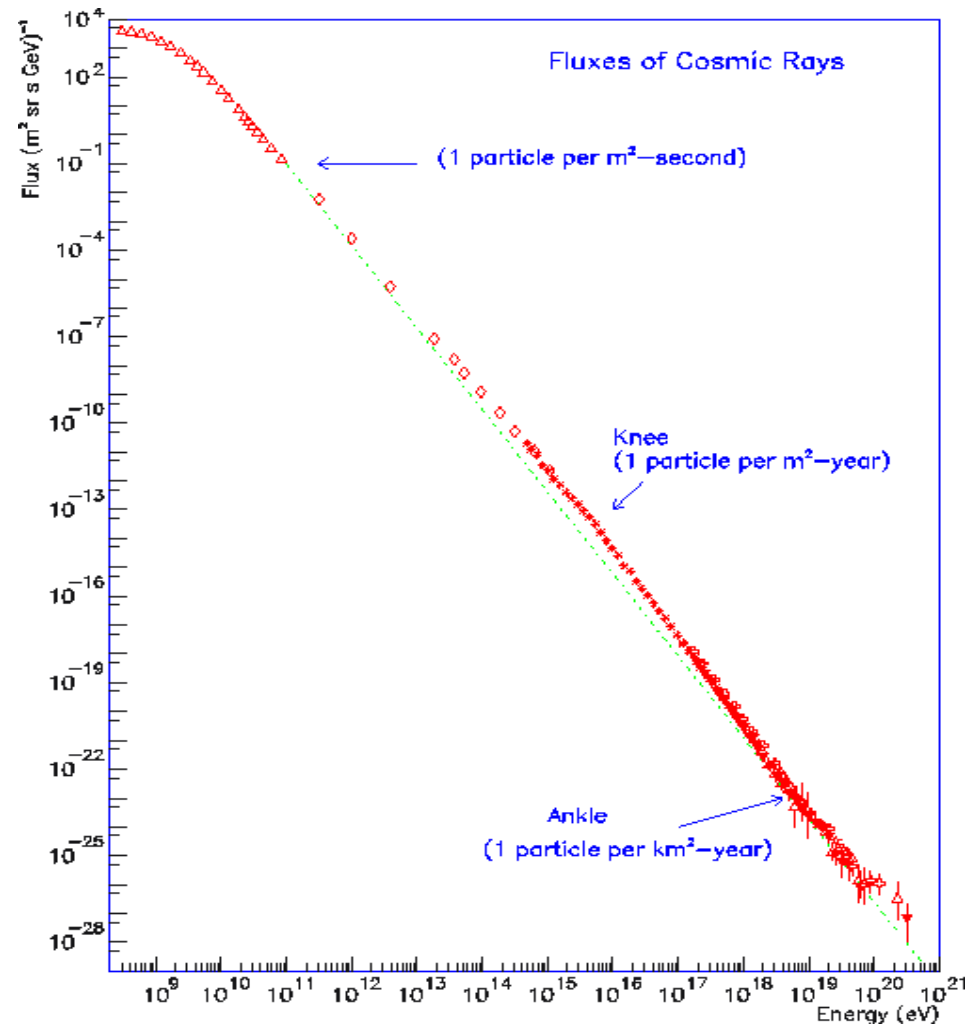


- Physics beyond the SM is expected at the EW scale $O(10^2 \text{ GeV})$: SUSY, Technicolor, Large Extra Dimensions, Little Higgs, ...
- SM extensions offer DM candidates

The quest for higher energies...



The Cosmic Ray beam at Earth



Ultra High Energy Cosmic Rays (UHECR) probe energies far beyond the reach of man made accelerators

UHECR

vs

LHC

Fixed target

$E \sim 10^{19} - 10^{20}$ eV

$\sqrt{s} \sim 100 - 400$ TeV

Low fluxes

Initial beam not well defined (p/Fe/
?)

Much poorer detection capabilities

Collider

$E \sim 10^{13}$ eV

$\sqrt{s} \sim 10$ TeV

High luminosities

Known beam composition

Powerful detectors

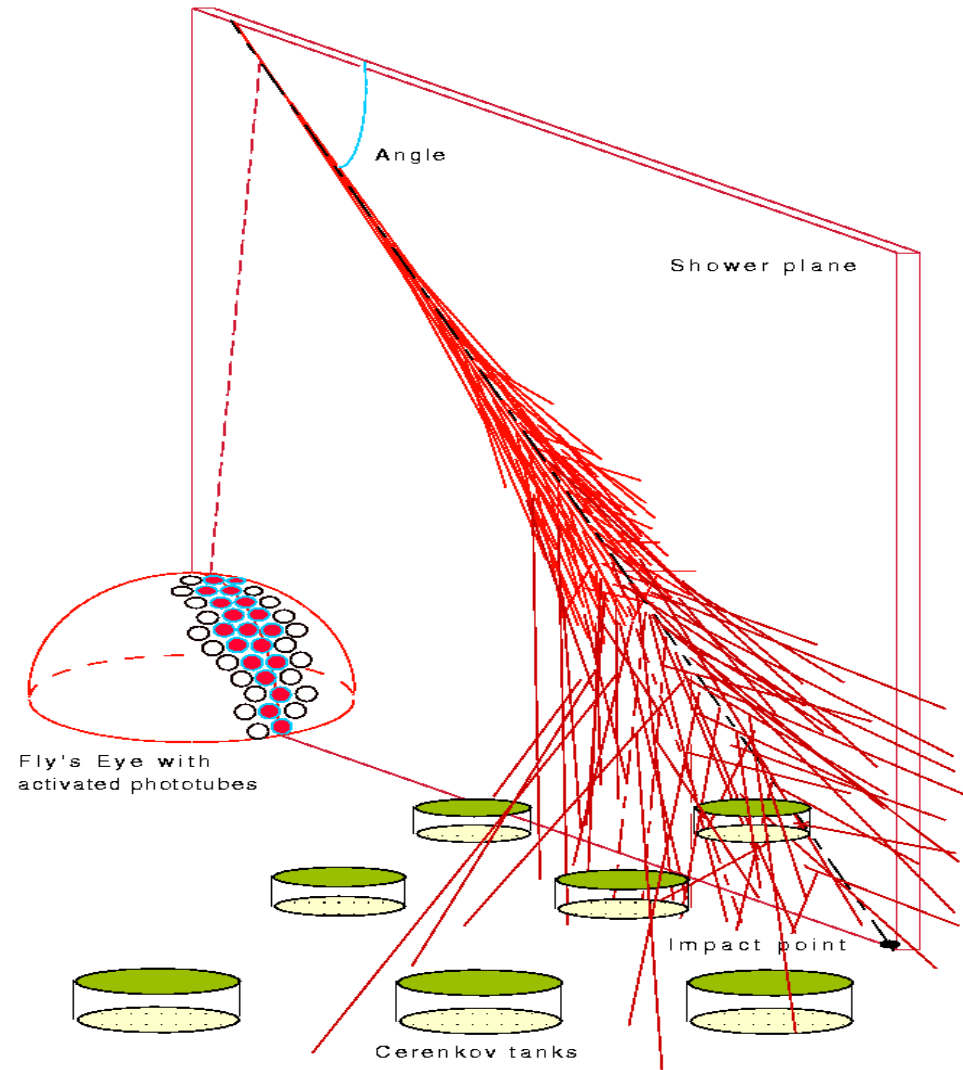
UHECR detectors can complement searches at man made accelerators

Exotics in UHECR

- Very large CM energies $O(10^2 \text{ TeV})$:
 - production of new heavy particles
- Huge beam energies/boost ($\gamma \sim 10^{11}$) :
 - Violation of fundamental physics laws ?
- Physics in extreme conditions:
 - look for signatures predicted by new theories/models but be prepared for the unexpected.

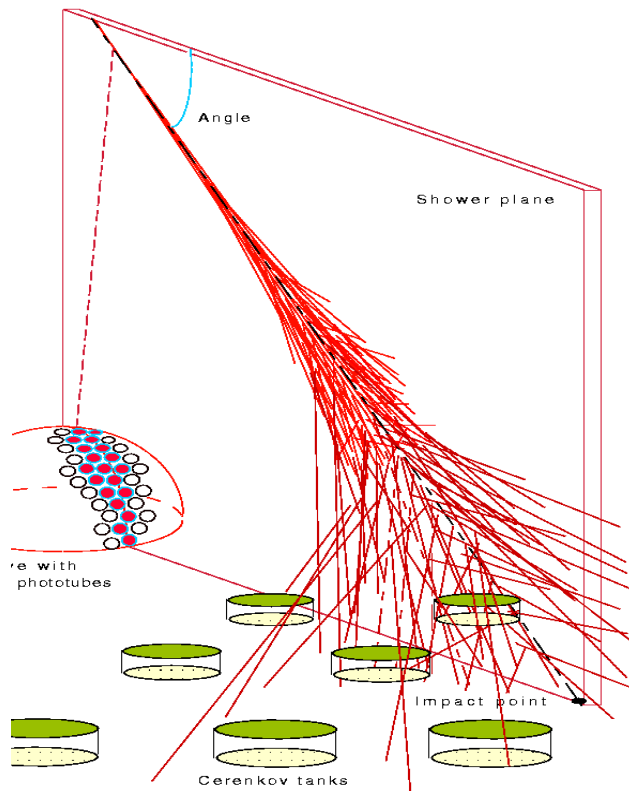
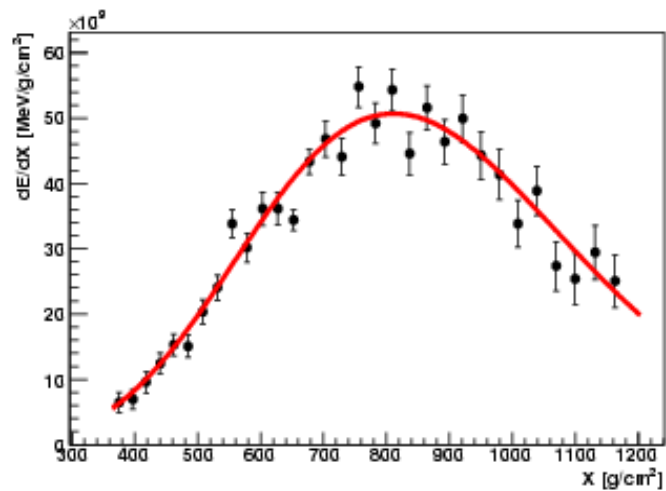
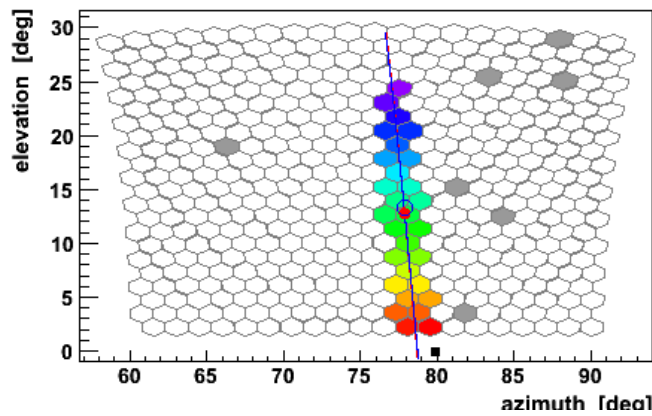
Measurement of UHE Cosmic Rays

Extensive Air Showers



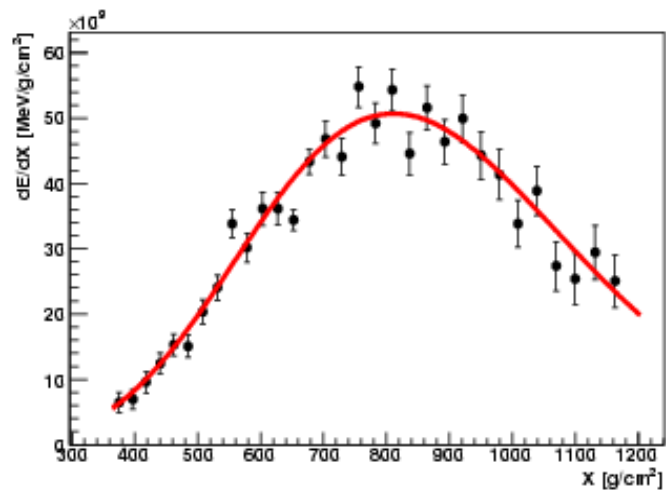
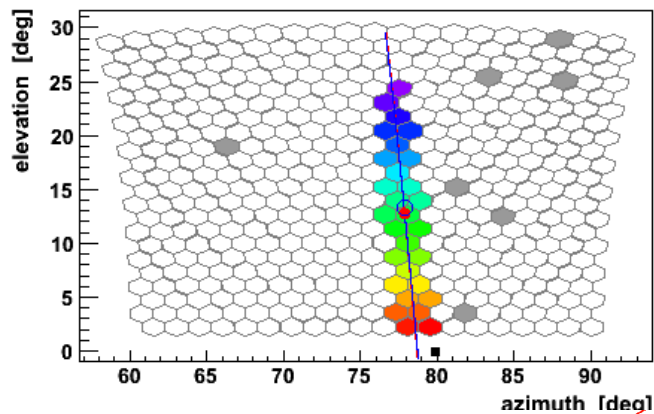
Measurement of UHECR

FD – Detection of Fluorescence from air N_2 molecules isotropic emission along the shower



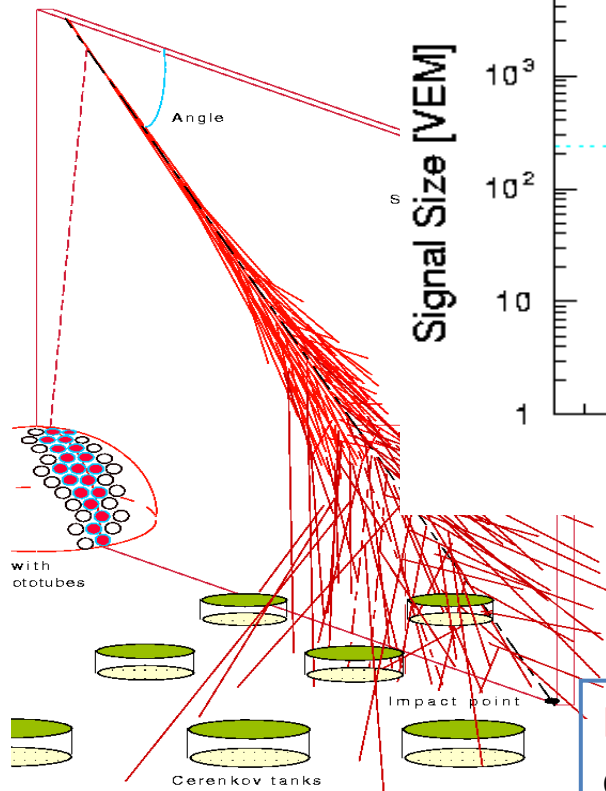
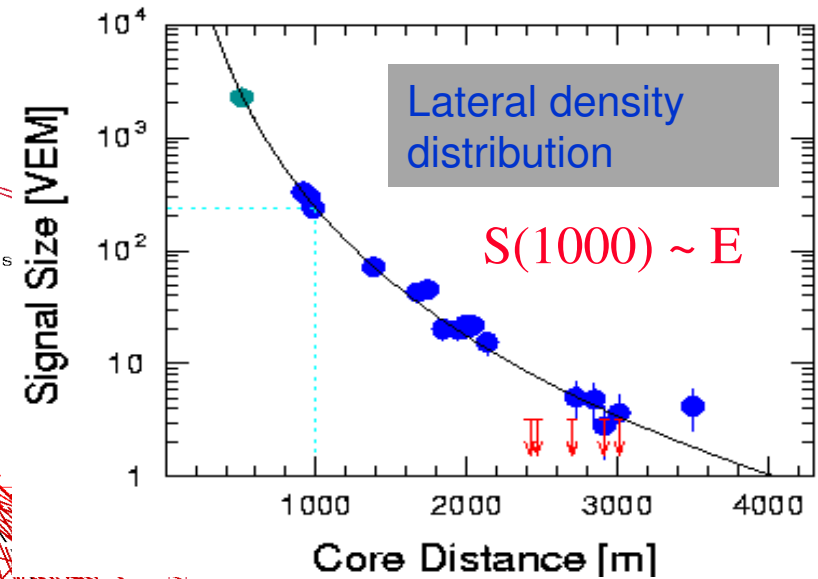
Measurement of UHECR

FD – Detection of Fluorescence from air N_2 molecules isotropic emission along the shower



SD-Direct detection of shower particles

ID 762238

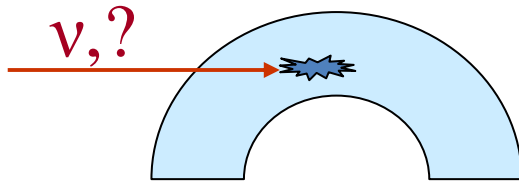


Hybrid Events - SD energy estimator calibrated by calorimetric measurement from Fluorescence Detector.

Physics Driven exotic targets

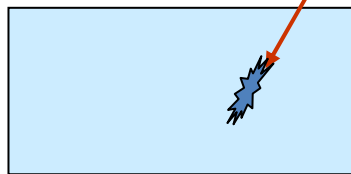
Identify specific exotic particles or interactions based on predictions in literature or searches by other experiments, or our own clever ideas.

- $\sigma_{em} < \sigma_{new} \ll \sigma_{had}$



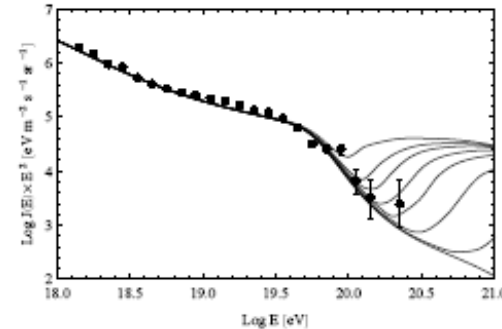
Mini-BH, l^* , LQ, χ^0 , ...

- $\sigma_{new} \sim \sigma_{had}$ p, ?



Z' , squarks, strangelets, Q-balls,
Exotic massive hadronic
particles, ...

- Lorentz invariance violation



Stecker
and
Scully

A modified GZK?

- Signatures

- modified profiles/double-bang
- Deep penetrating showers
- Shower speed
- Gamma showers, Muon contents
- ????

Topological searches ???

(M. Pimenta)

Exotic particle searches at UHECR

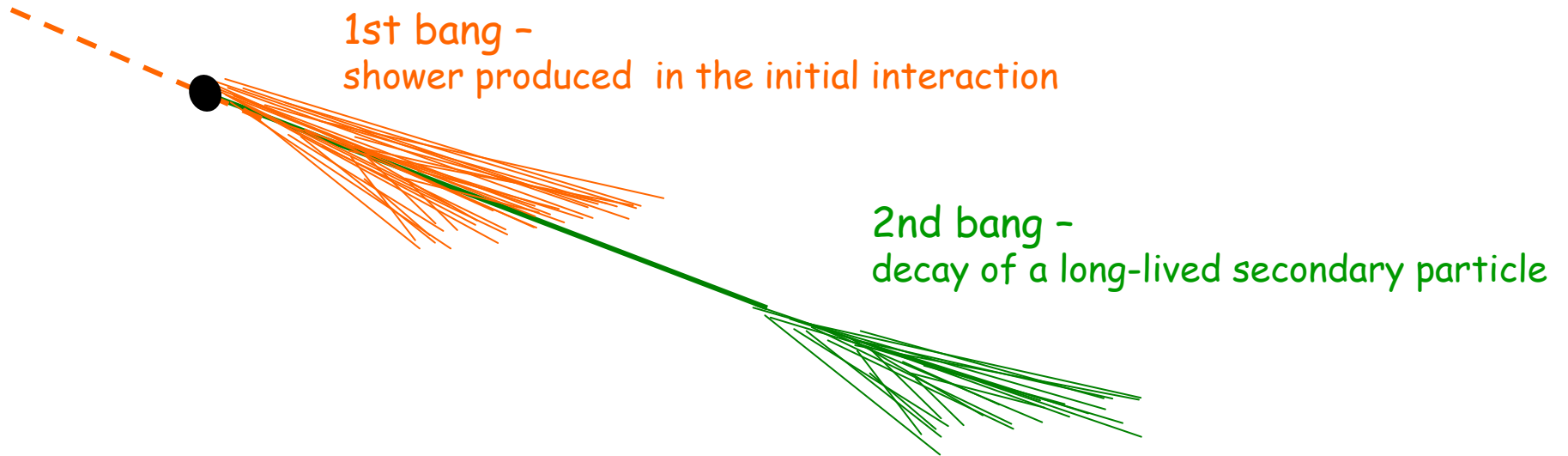
| Candidate Particle | Properties | Observables |
|------------------------|---|--|
| Magnetic Monopole | Mass of about 10^8 GeV, created in $SU(4) \times SU(2) \times SU(2)$ symmetry breaking, catalyzes proton and electron decay | Strong electromagnetic component in the form of large amounts of Cherenkov radiation (~8500x more than a muon), shallow X_{max} |
| Strangelet | Theorized ground state of QCD, bags of roughly equal proportions of up, down and strange quarks, mass of $10^4 - 10^8$ GeV | Interacting primarily hadronically, narrow Gaisser-Hillas with shallow X_{max} |
| Micro-Black Hole | The amount of Hawking radiation produced by a black hole is inversely proportional to its mass, so a burst of hadrons and photons would be expected from an MBH | Since secondary MBH's from neutrino-nucleon interaction is most likely looking for hadronic bursts with very high zenith angle (>75 degrees) is most promising |
| Q-ball | Created by Bose-Einstein condensation of a SUSY strangelet in which the fermionic quarks become bosons and then relax to a single wavefunction | Proton decay catalysis from electrically neutral Q-balls creating largely electromagnetic, deeply penetrating showers |
| Other Exotic Heavy | Generic, high mass, low rigidity particle, possibly a relic particle or dark matter candidate | Without saying anything else about the properties of the particle, the most telling observable would be its velocity being less than c |
| Other Strange Profiles | Short lived secondaries being produced from an initial interaction (taus), rapidly decaying exotic in the atmosphere, other strange or unexpected behavior | Double Bangs, Bumpy showers, FD-SD mismatches, unusually strong signal in a single SD with low or no other signal |

(D. Schuster)

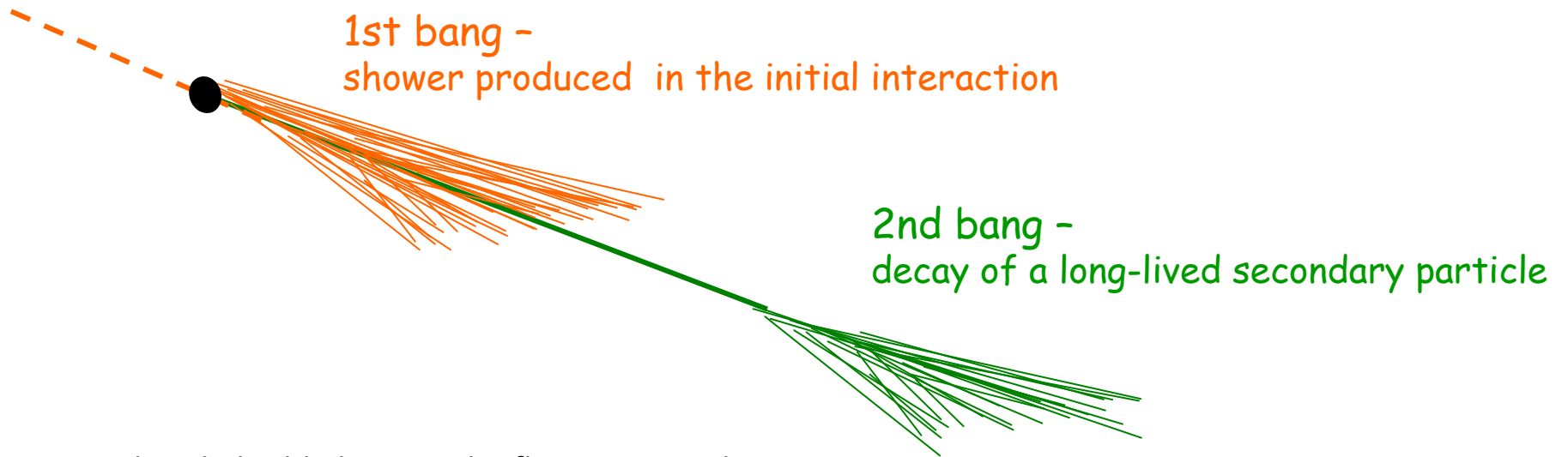
Signatures:

Modified shower profile, depth of shower max, shower speed, muon content, Cherenkov radiation, ...

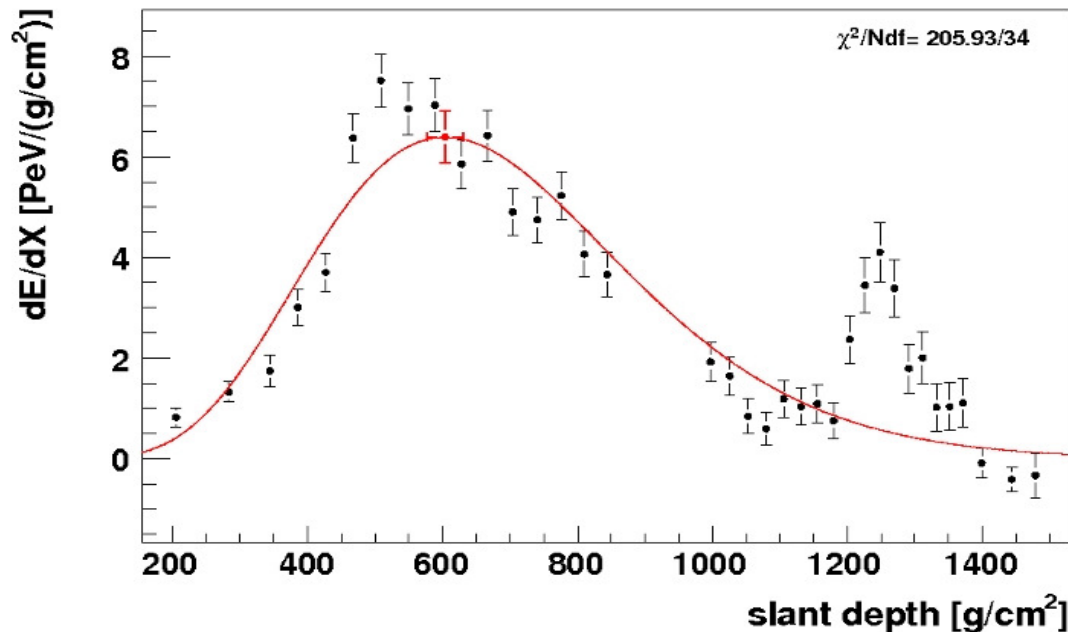
The double bang signature



The double bang signature



Simulated double bang in the fluorescence detector

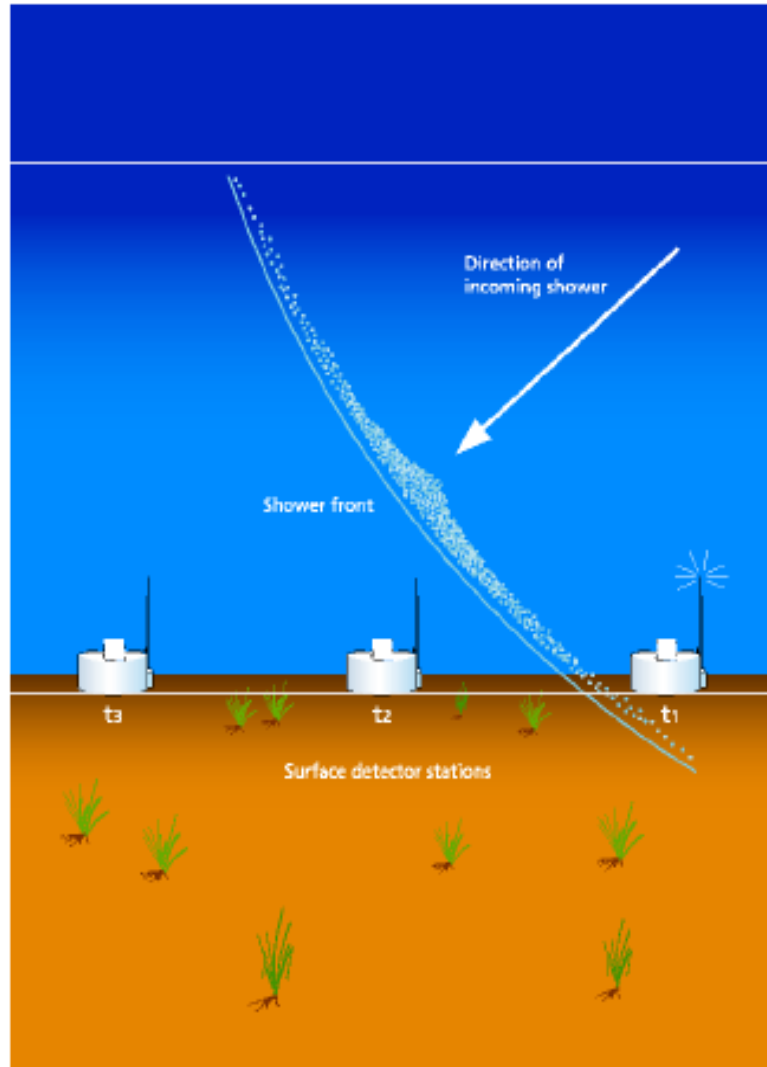


Clear double bang, but...

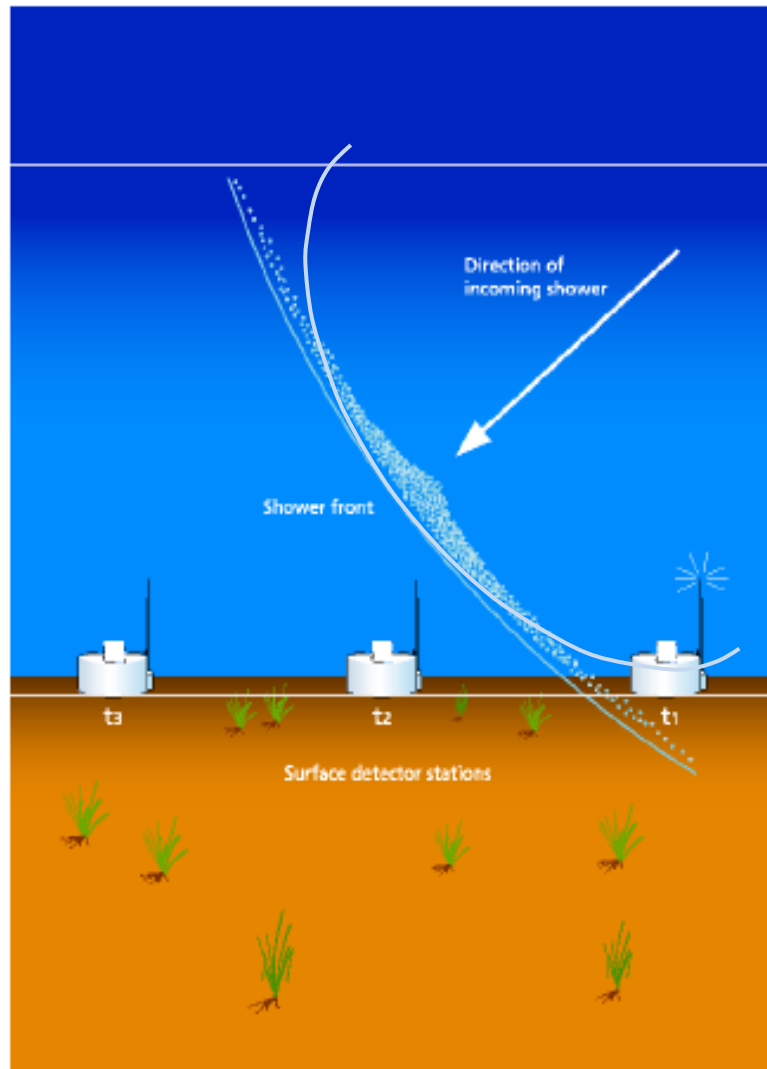
Signature might be less distinctive
(e.g. only a distortion of the
shower shape)

Atmospheric conditions (clouds,...)
may give rise to bumpy showers

Double shells in the surface detectors



Double shells in the surface detectors



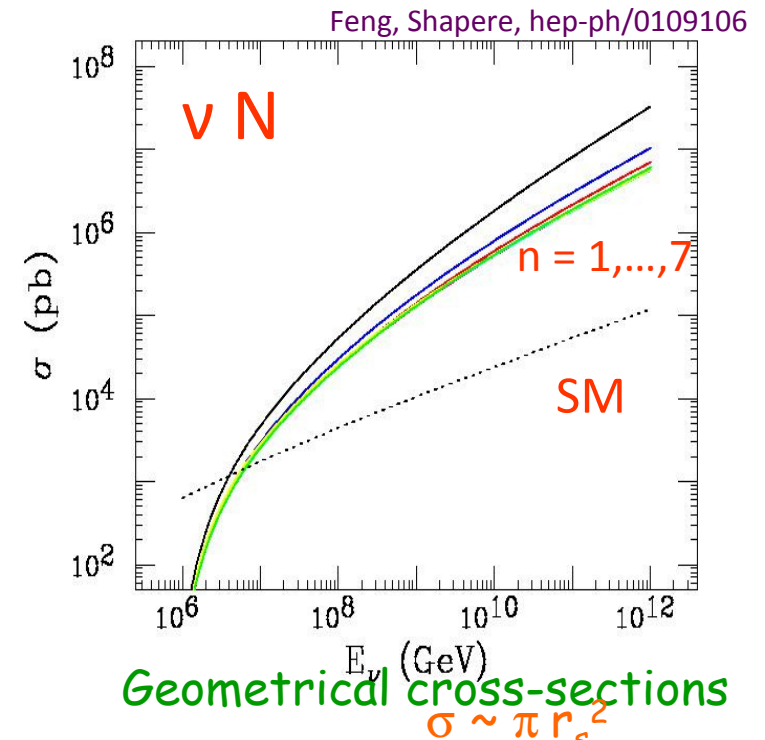
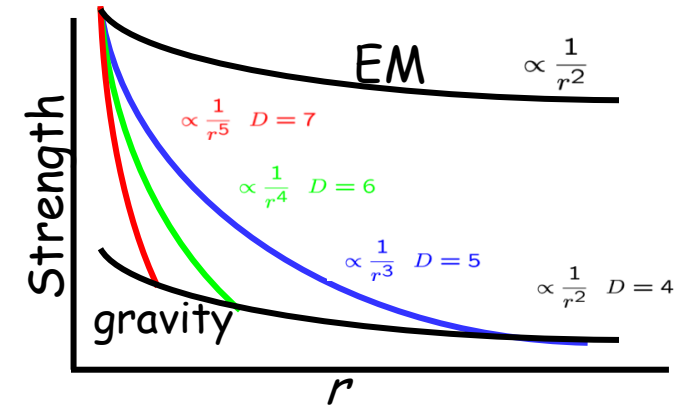
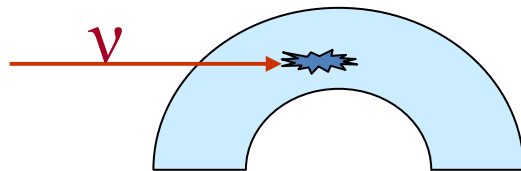
- Correlated peaks in different stations of one same event
- Analyze rare signal structures and measure the degree of correlation

Exotics in UHECR

- TeV black holes
- UHECRONS
- Lorentz invariance violation

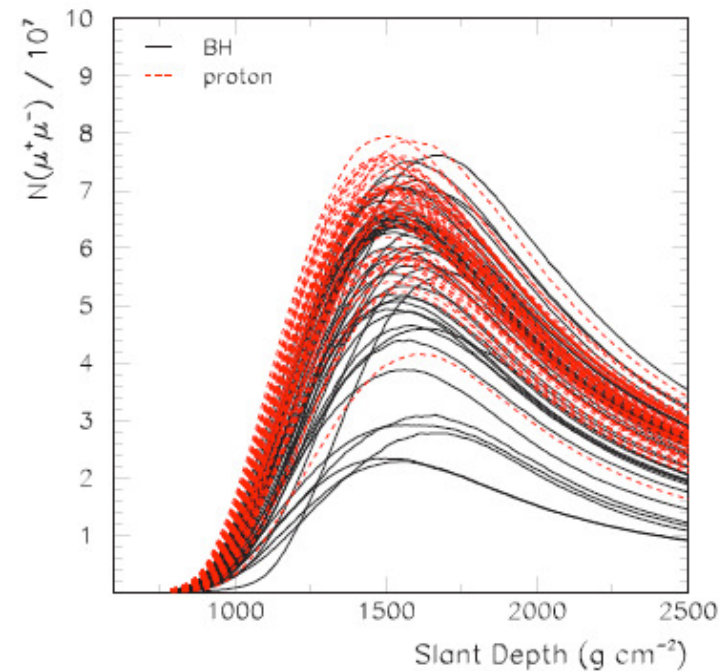
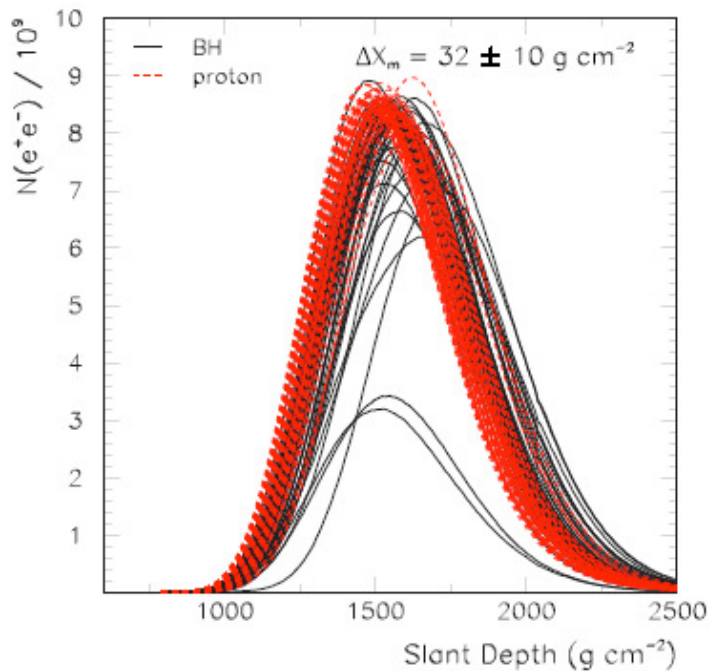
TeV Black Holes in UHECR interactions

- Planck scale at $O(\text{TeV})$ in scenarios with large extra dimensions
- Gravity gets stronger at shorter distances
- Black holes can be produced in collisions with $\sqrt{s} > \text{TeV}$
- Mini BH production in UHECR : $\nu + N \rightarrow \text{BH}$



Mini BH detection

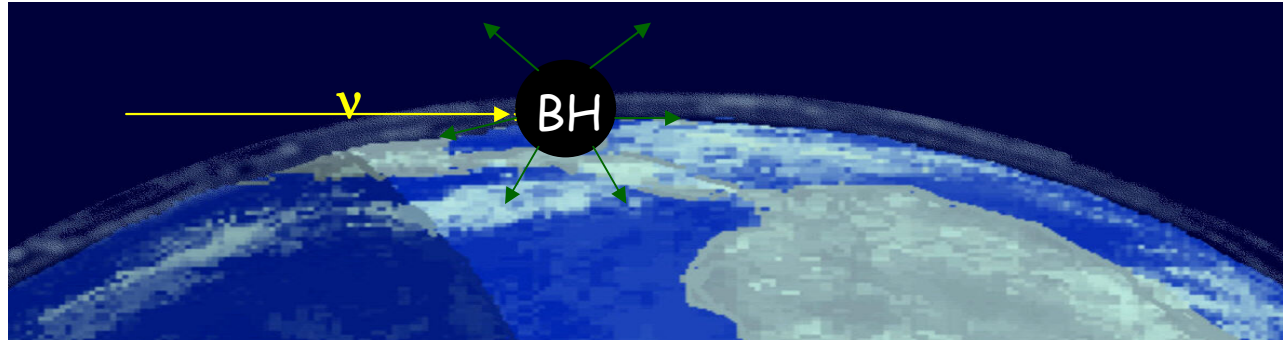
- Main features of BH showers are independent of details of BH formation and evolution (Ahn & Cavaglia - PRD 73 - 2006)



- Hadronic showers with large fluctuations; not so different from proton showers
- Experimentally tough...

Mini BH detection: the double bang scenario

V. Cardoso, M.C. Espirito Santo, M. Paulos, M. Pimenta, B. Tomé
Astroparticle Physics 22 (2005)



- Horizontal showers starting deep in the atmosphere
- 1st bang : Mini BH production + instantaneous democratic decay
- 2nd bang : decay of energetic taus ($L_{\text{int}} \gg L_{\text{dec}}$)
- Distinctive signature ! But a narrow window for the detection of the second bang (tau energy vs tau decay length)

Next generation space-borne telescopes should take it into consideration

Observation prospects

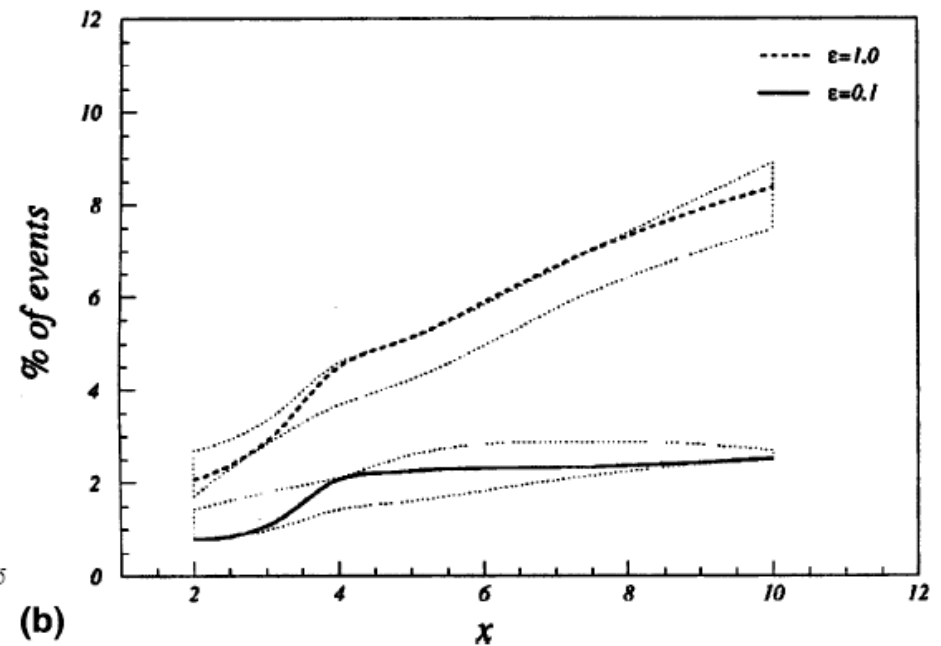
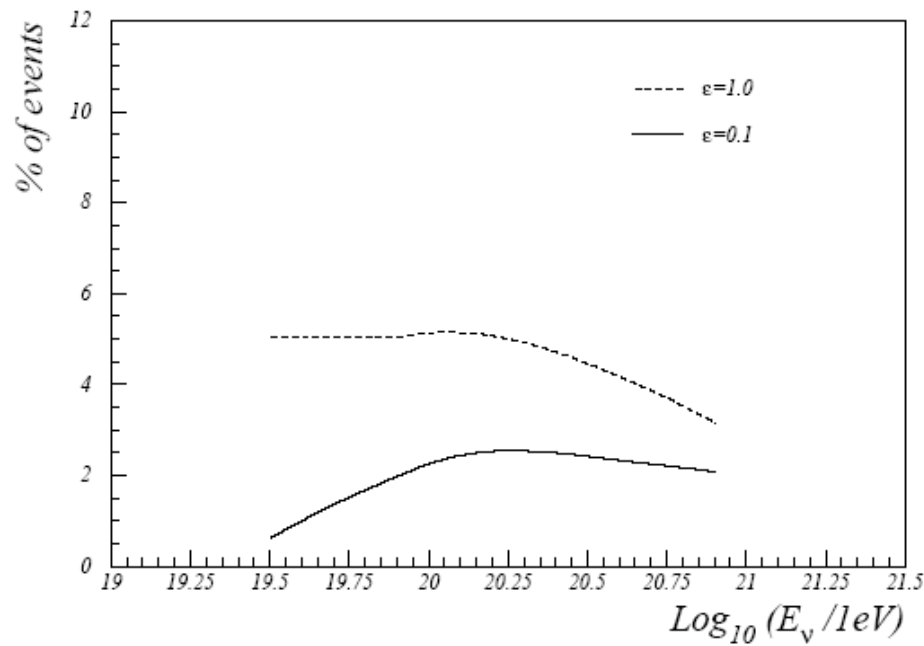
EUSO taken as case study

Order of magnitude computation with simple model

GIL + fluo model (Kakimoto) + LOWTRAN for atm

First bang randomly distributed horizontally, and vertically according to density profile

Chances of seeing the 2nd bang once the 1st one is detected



A few % of the events give double bangs in EUSO

UHECRONS

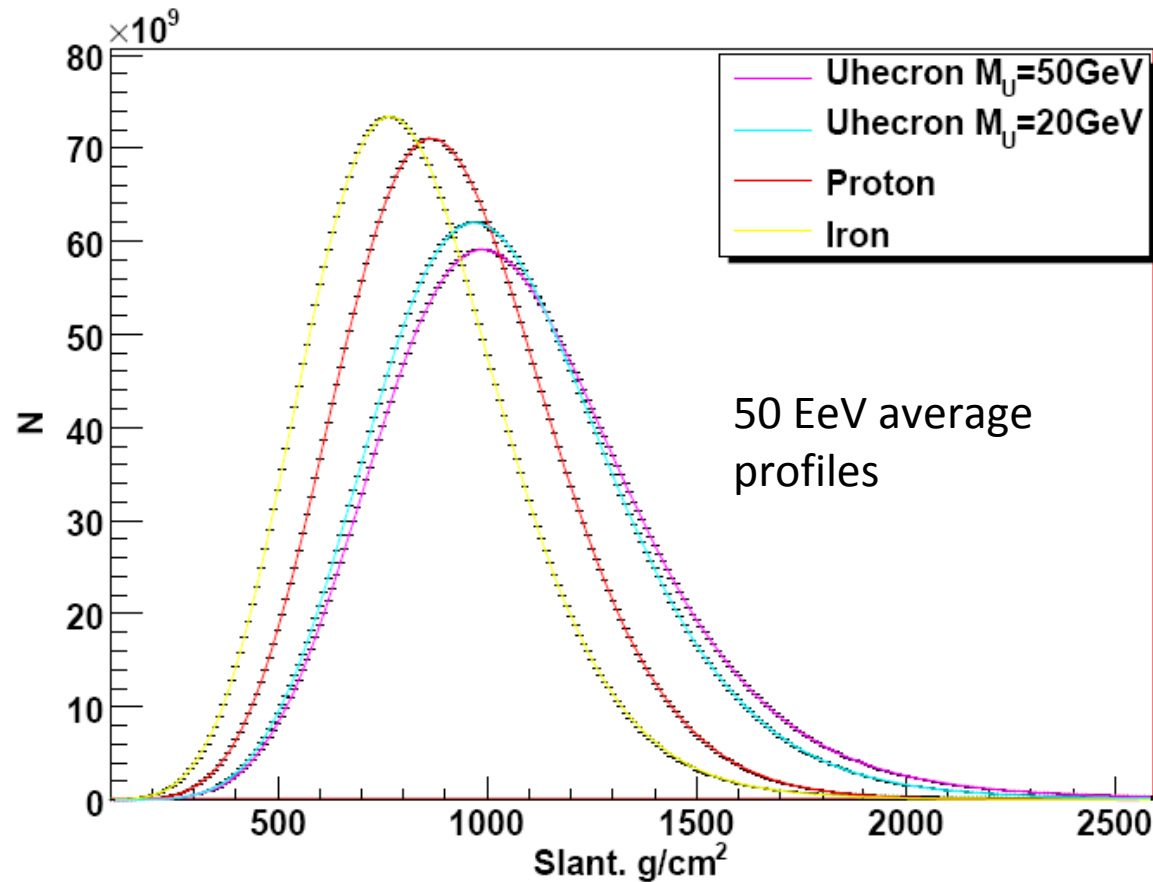
Chung, Farrar and Kolb Phys. Rev. D 57 (1998)

- Exotic massive stable neutral hadrons
- Increased threshold energy for pion production -> sources beyond GZK distances
- Candidates:
 - Glauino-gluon color singlet in SUSY models in which the gluino is the LSP (S.Raby)
 - Strongly interacting WIMPless dark matter (J. Feng)

UHECRON detection

Washington Carvalho Jr. & Ivone Albuquerque, to be published in PRD

Main distinctive signature is the longitudinal profile of EAS

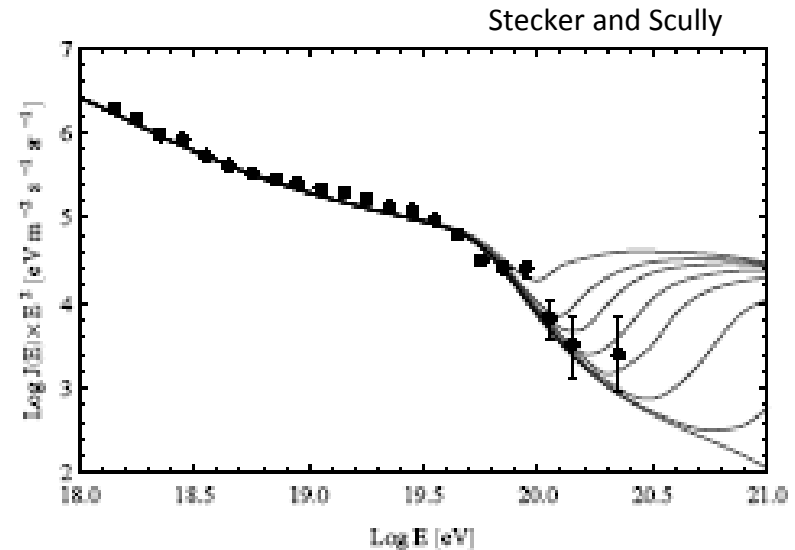


- Slower shower development
- Deep maximum
- Surface detector information will improve UHECRON discrimination (different $\mu/e.m.$ ratio !!)

UHECRON with $M < 50$ GeV can be detected and discriminated against p, Fe

Lorentz invariance violation

- Photoproduction $\gamma p \rightarrow \pi p$ at $\gamma \sim 10^{11}$
- Different propagation through the CMB ?
- Modified GZK cutoff



- In certain conditions no π^0 decay above given threshold energy
- Modified shower development (shower shape, muon content)

New physics observed in standard UHECR collisions !

The Muon Energy Spectrum as a Test of the Validity of Special Relativity at Small Distances.

M. DARDO, G. NAVARRA and P. PENENGO

Laboratorio di Cosmo-Geofisica del CNR
Istituto di Fisica Generale dell'Università - Torino

(ricevuto il 12 Dicembre 1968)

Summary. — The recent observation of Lundberg and Rédei that a violation of the principles of special relativity at small distances would cause a change in the lifetime of pions and muons is submitted to a test against cosmic-ray data. It is shown that at sufficiently high energies the muon energy spectrum and the zenith angle distribution of muons will be affected by the assumed violation. Present experimental data permit us to establish an upper limit of $5 \cdot 10^{-17}$ cm for the dimensions in which special relativity may not be valid.

Summary

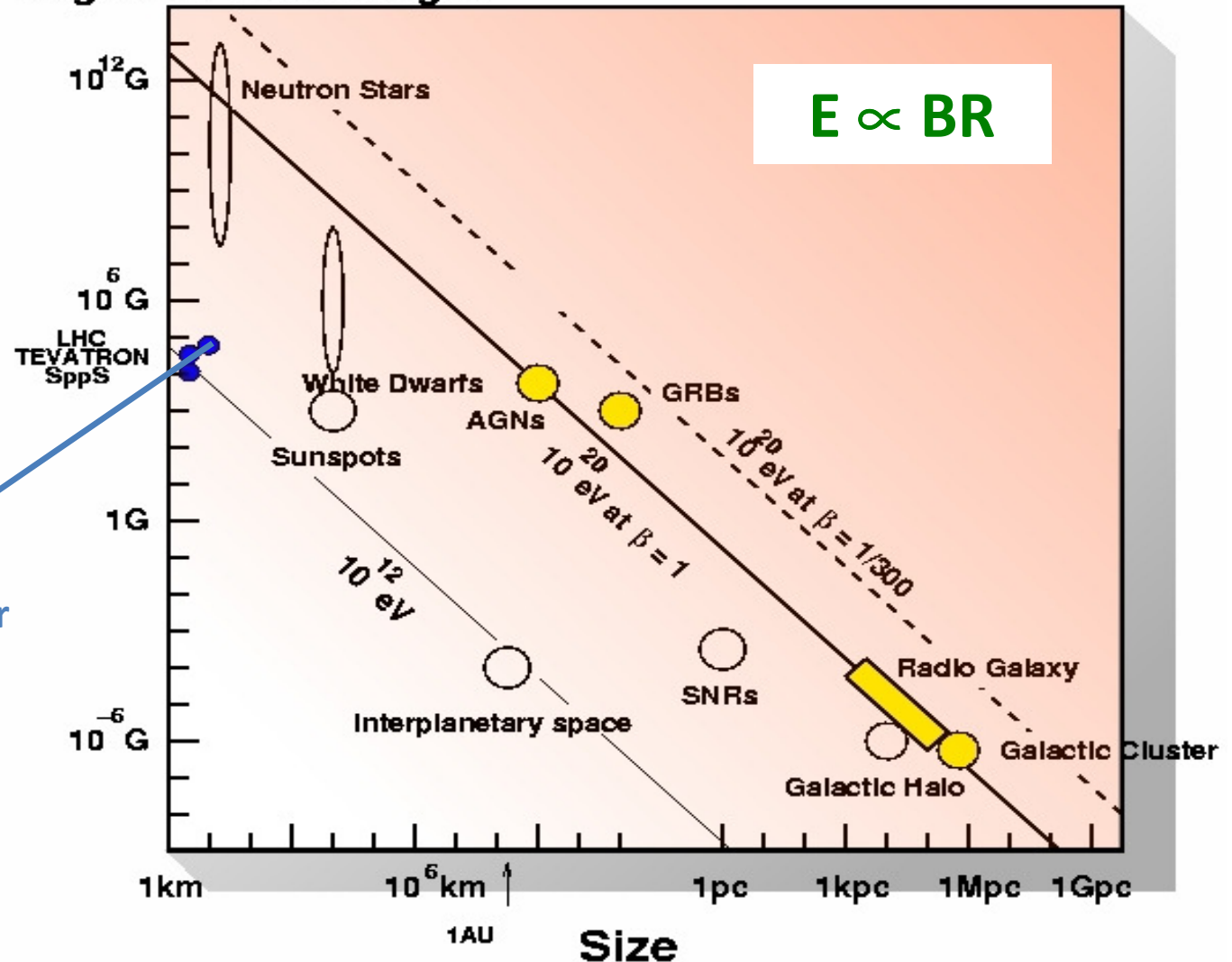
- UHECR telescopes probe energies far beyond man made accelerators
- Exotic physics is probably hiding in the bulk of the cosmic ray events
- Be prepared for the unexpected !

BACKUP

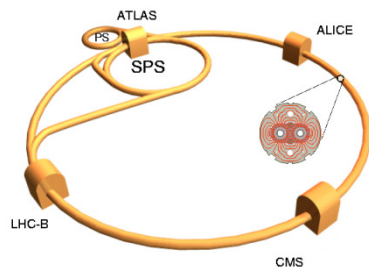
Cosmic accelerators

Hillas-plot

Magnetic Field Strength

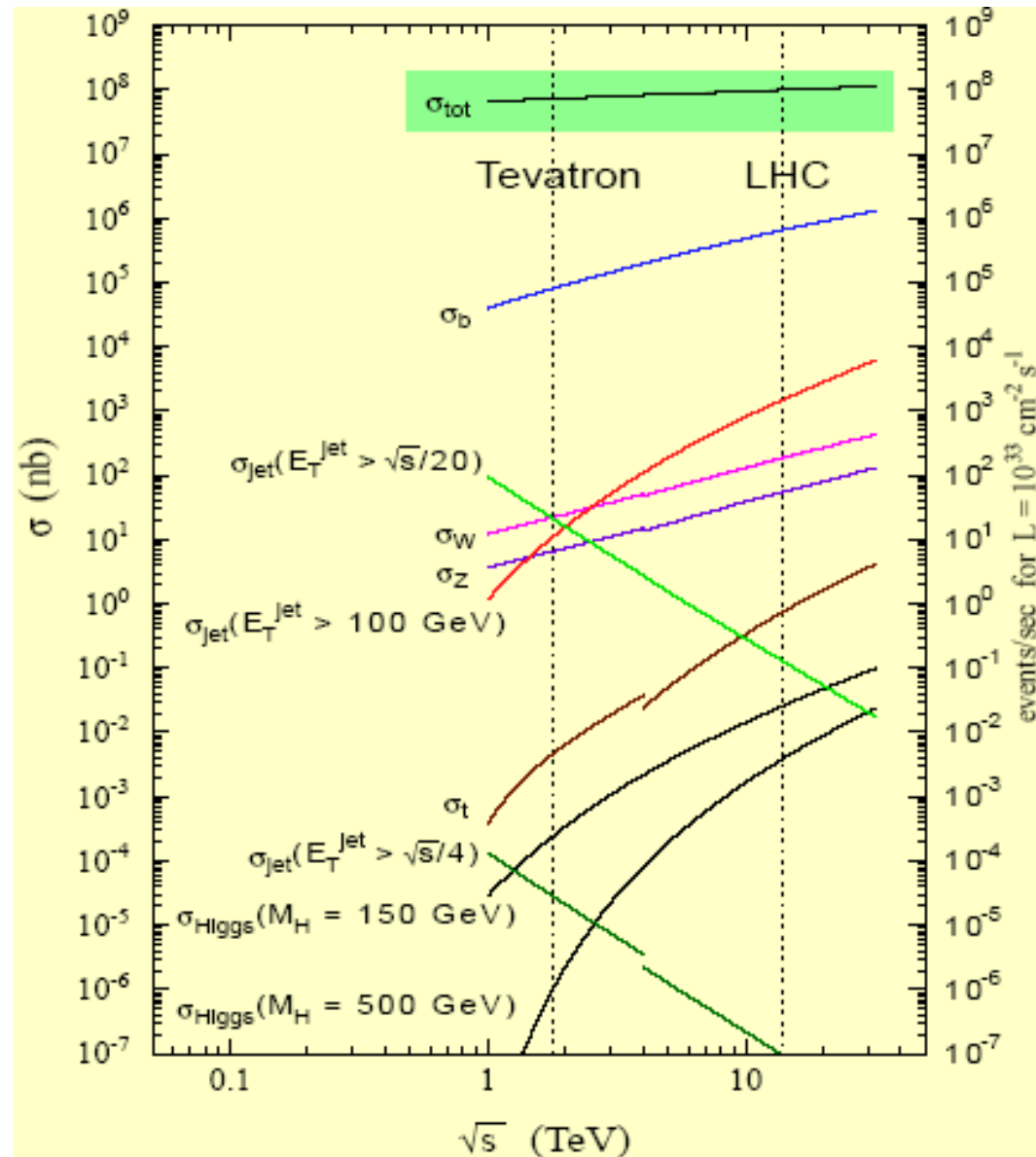


Large Hadron Collider



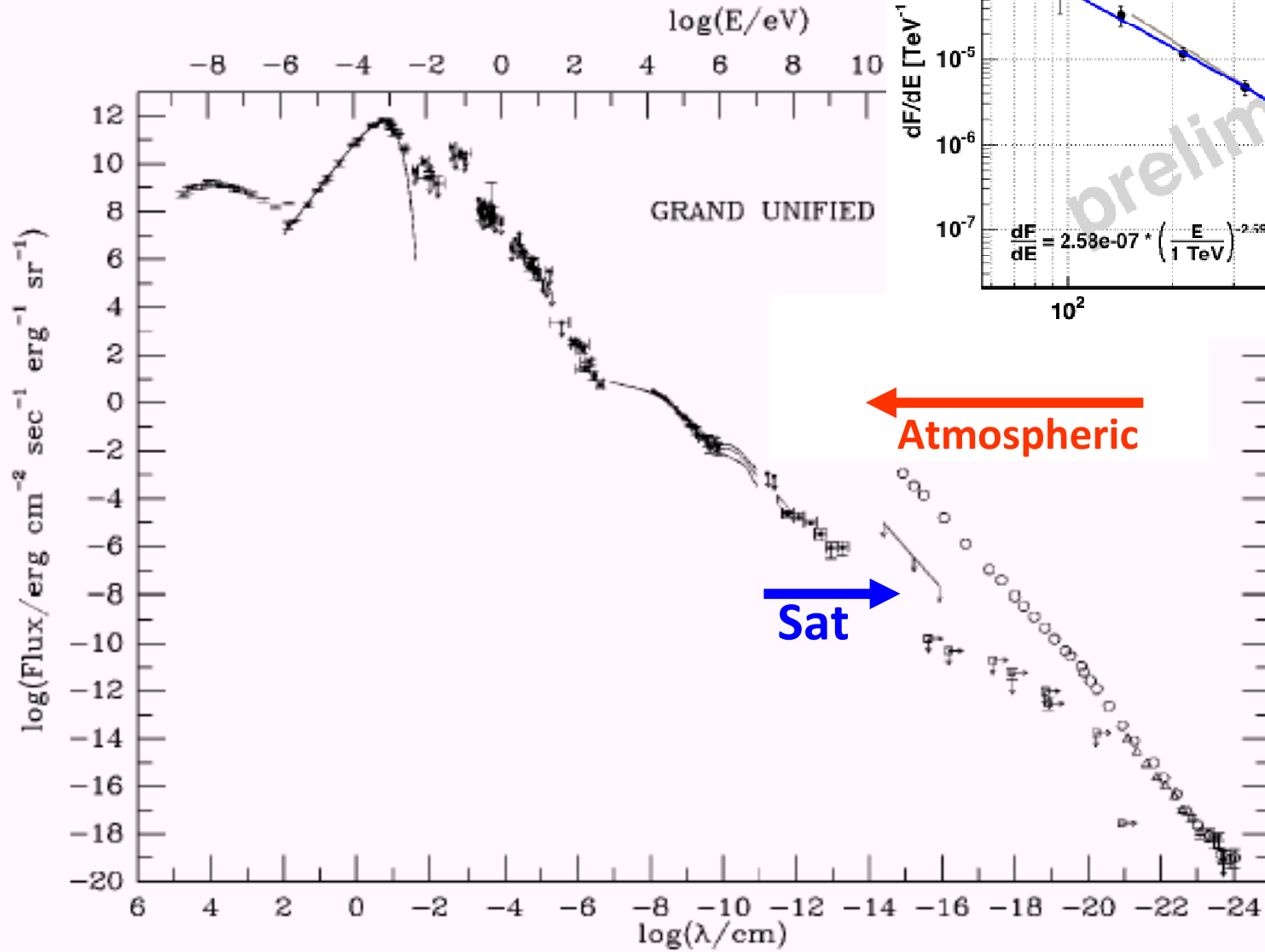
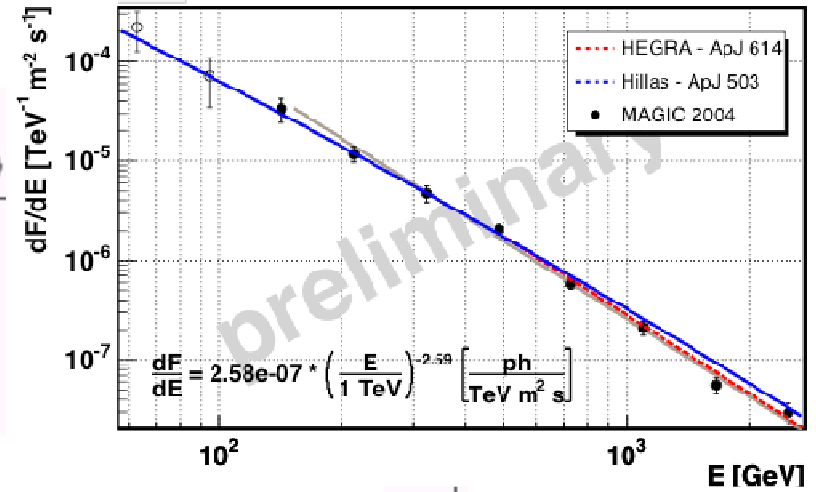
$R \sim 10$ km, $B \sim 10$ T
 $E \sim 10$ TeV

pp SM background



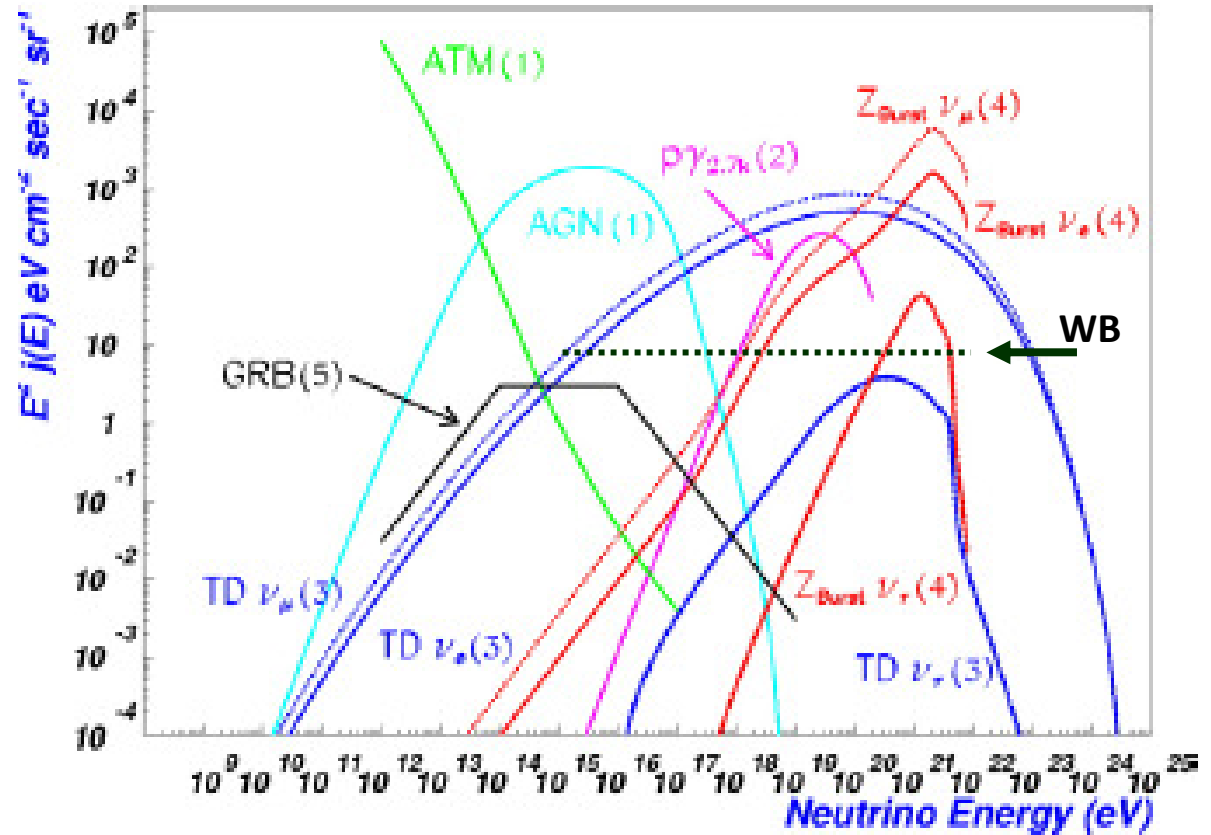
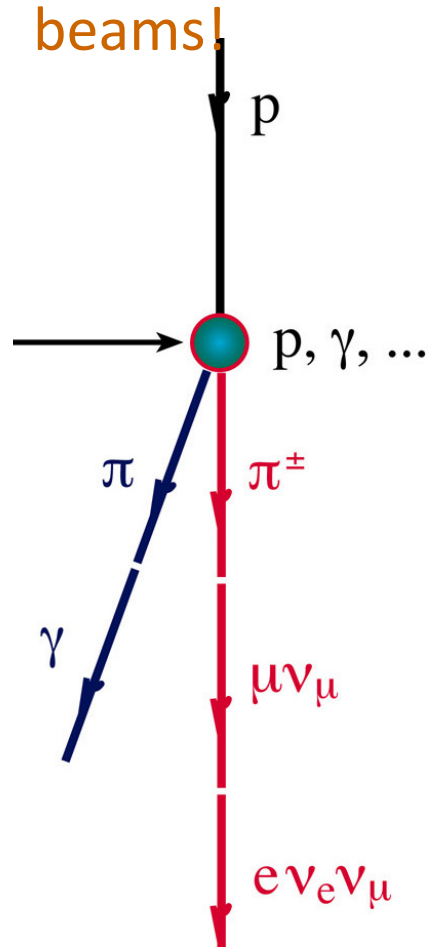
γ beams

 **MAGIC** Crab energy spectrum

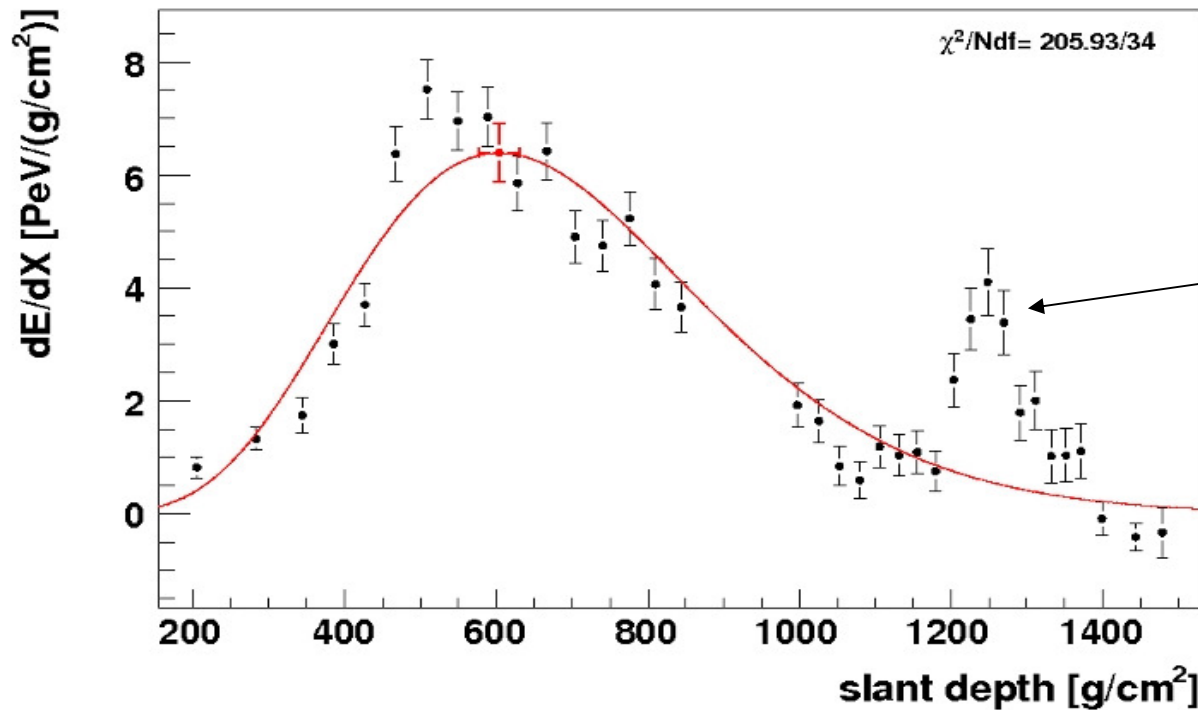


ν beams

secondary
 γ and ν
beams!



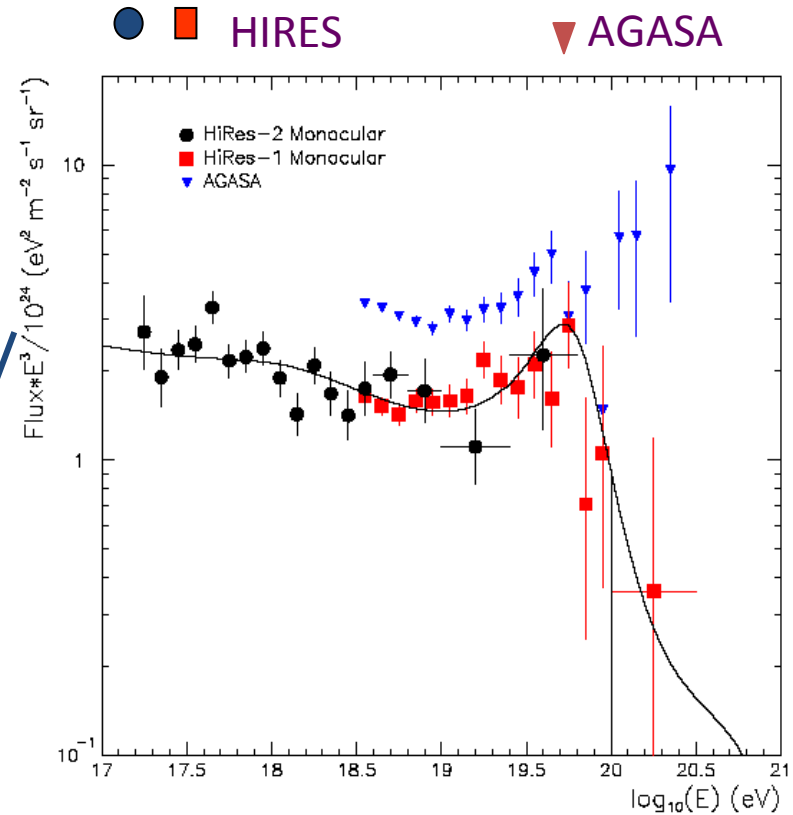
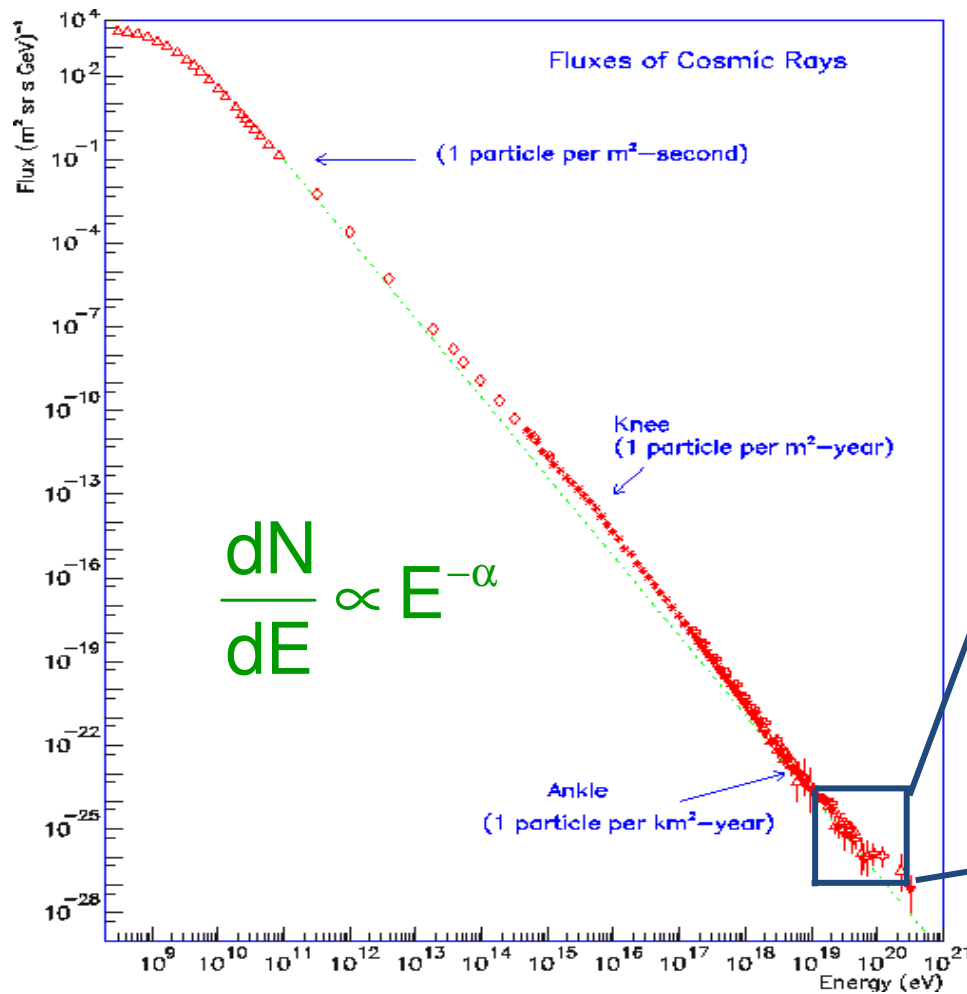
BUT no measurements
above $E > 10^6$ GeV !



This analysis not sensitive to something like a double bang (profile caused by poor atmospheric conditions)

There are two ways of doing an exotic search, one focuses on finding strange things (like double bangs) that defy the typical profile. The other looks for things that are similar to typical profiles, but slightly differ. This analysis is sensitive to the latter and may help answer the questions: “how sensitive is our detector to exotic things?” and “how much room is in our data for exotic things?”

Charged (p/nucleus) beams



~2001

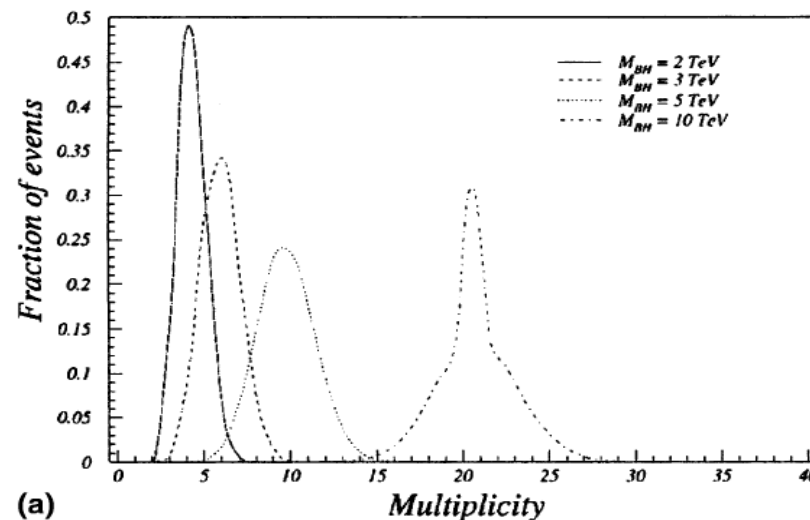
$E > 10^{20}$ eV ~ 20 events

The first bang: production & decay of BH (I)

Mini-BH production and decay simulated with CHARYBDIS

Harrison, Richardson, Webber, hep-ph/0307305

- Developed to simulate BH production in hadron colliders
- Careful treatment of BH decay. Pythia for standard particle decay and hadronisation
- Democratic decay into all SM particles
- Conservation of charge and total angular momentum. No lepton number conservation



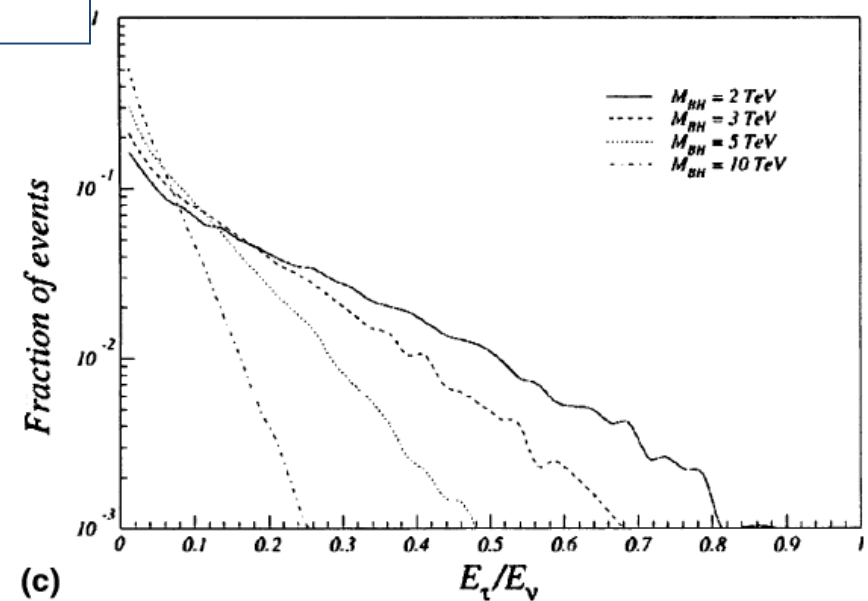
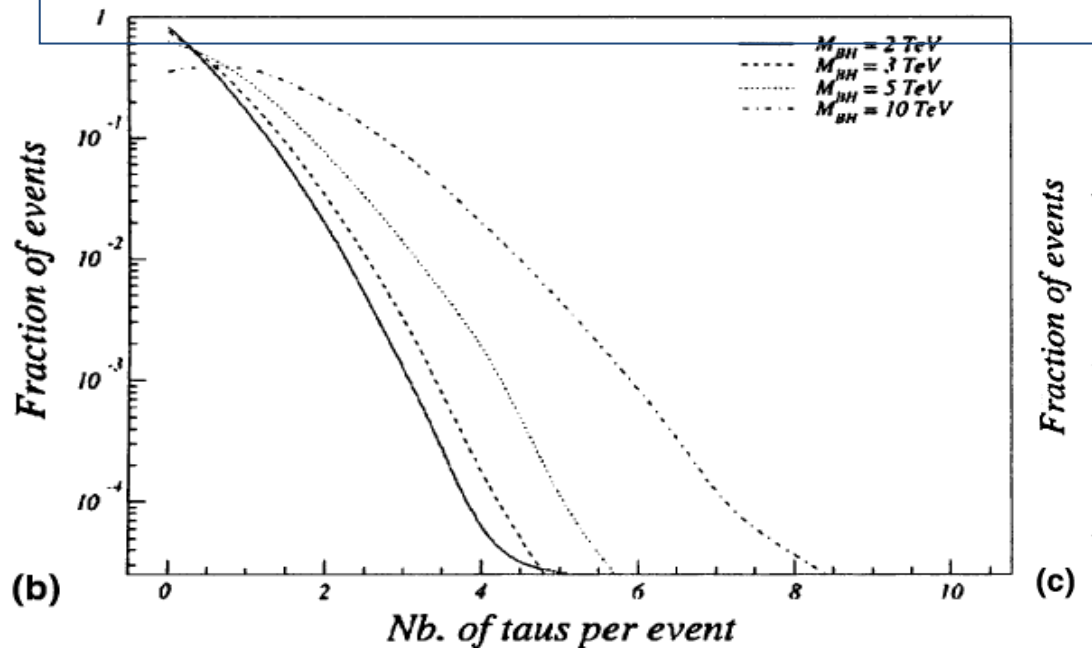
$M_D = 1 \text{ TeV}, n=3$

(a)

The first bang: production & decay of BH (II)

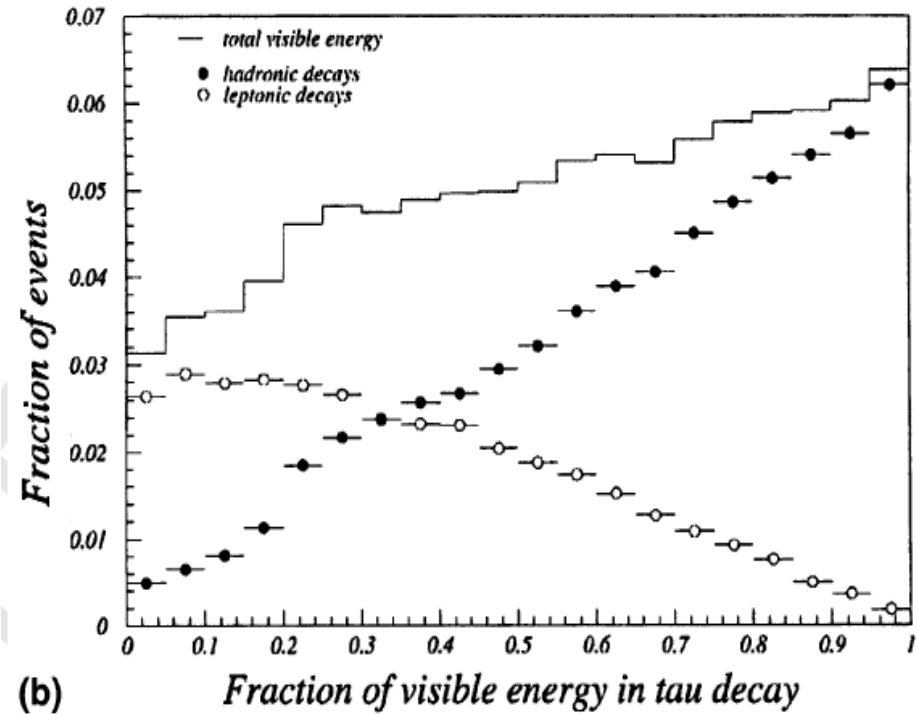
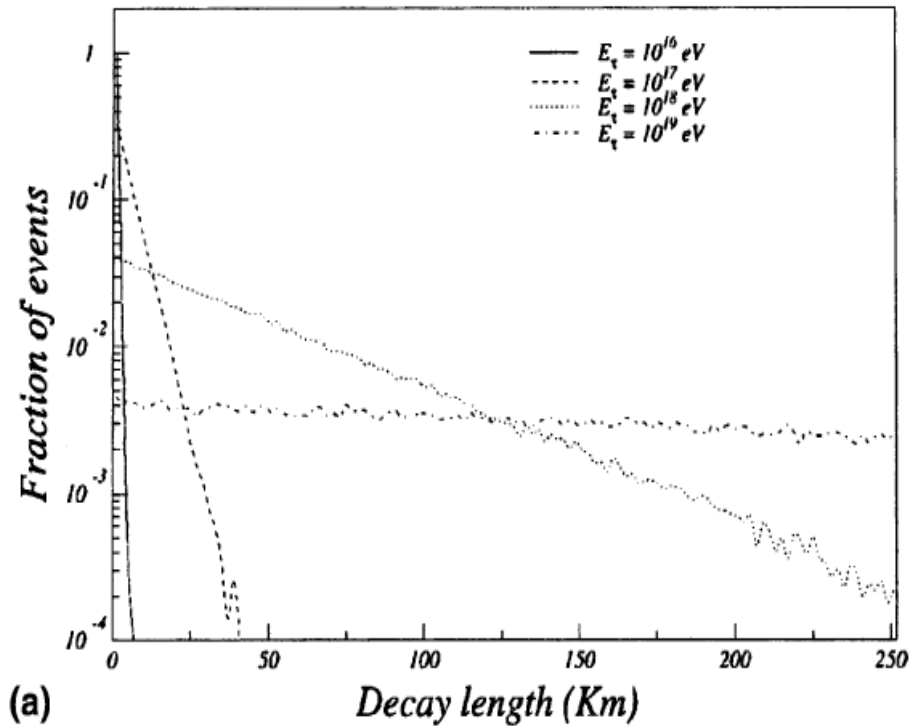
- Settings: nb of extra dimensions $n=3$ and $n=6$
 Planck mass $M_D=1,2,3$ TeV
 fixed BH masses $M_{BH}=2M_D, 3M_D, 5M_D$
- No explicit cross-section model. Fixed BH masses.
 Results should be convoluted with a given M_{BH} spectrum
 (dependent on pdf extrapolation and specific cross-section model)
- Main parameter is $x=M_{BH}/M_D$. Dependence on n is moderate

| MBH/TeV | $\langle N\tau \rangle$ |
|---------|-------------------------|
| 2 | 0.2 |
| 3 | 0.3 |
| 5 | 0.5 |
| 10 | 1.1 |



The second bang: the decay of energetic taus

- Taus from bh decays and W/Z/t decay treated using pythia
- $L_{\text{int}} \gg L_{\text{dec}}$



Observation prospects (I)

Approach:

EUSO taken as case study

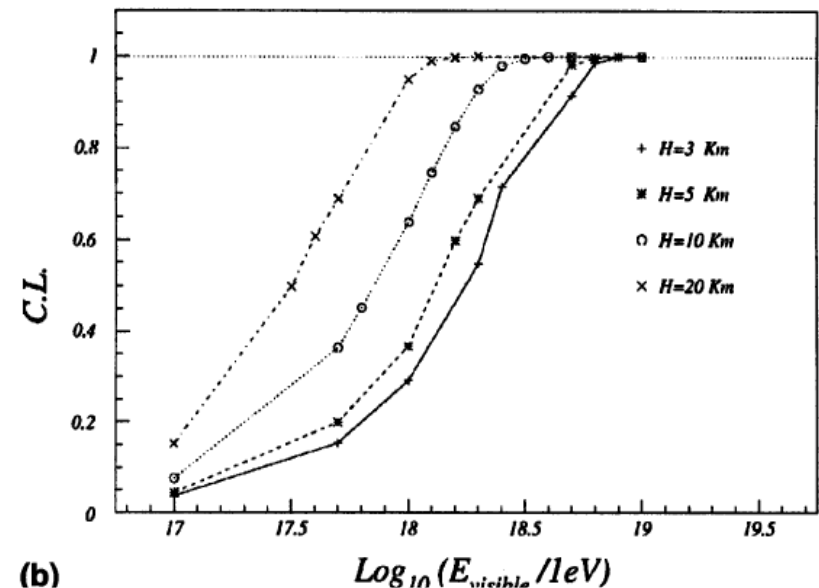
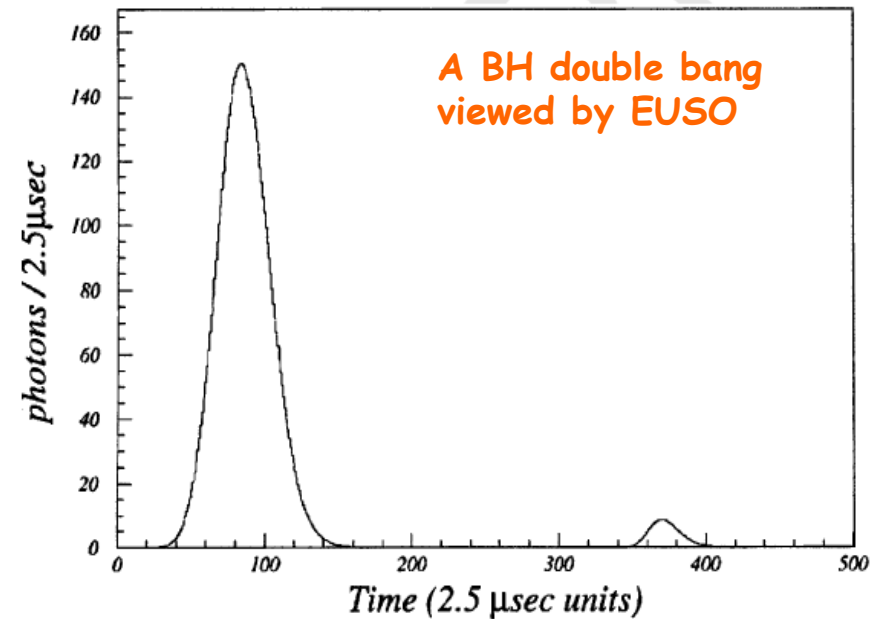
Order of magnitude computation with simple model
GIL + fluo model (Kakimoto) + LOWTRAN for atm

First bang randomly distributed horizontally, and vertically according to density profile

Observation window constrained by geometry (FOV) and signal to noise ratio

Lowering the energy threshold for the 2nd bang:

- Aligned with 1st bang in the focal surface
- Modified Frequentist likelihood ratio method:
Taking into account signal & background shape



UHECRON @ 100 EeV

